

2017 UCTA Summer Student Program

- (1) Digging in the ALMA data archive
- (2) Searching for young proto-planetary disks from ALMA archival data

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outline

introduction

Example
(analysis)

conclusion

Keplerian disk

- ▶ We use moment map to study every disk ,if the disk obeys kepler's law ,we call it keplerian disk .It obey's the following equation

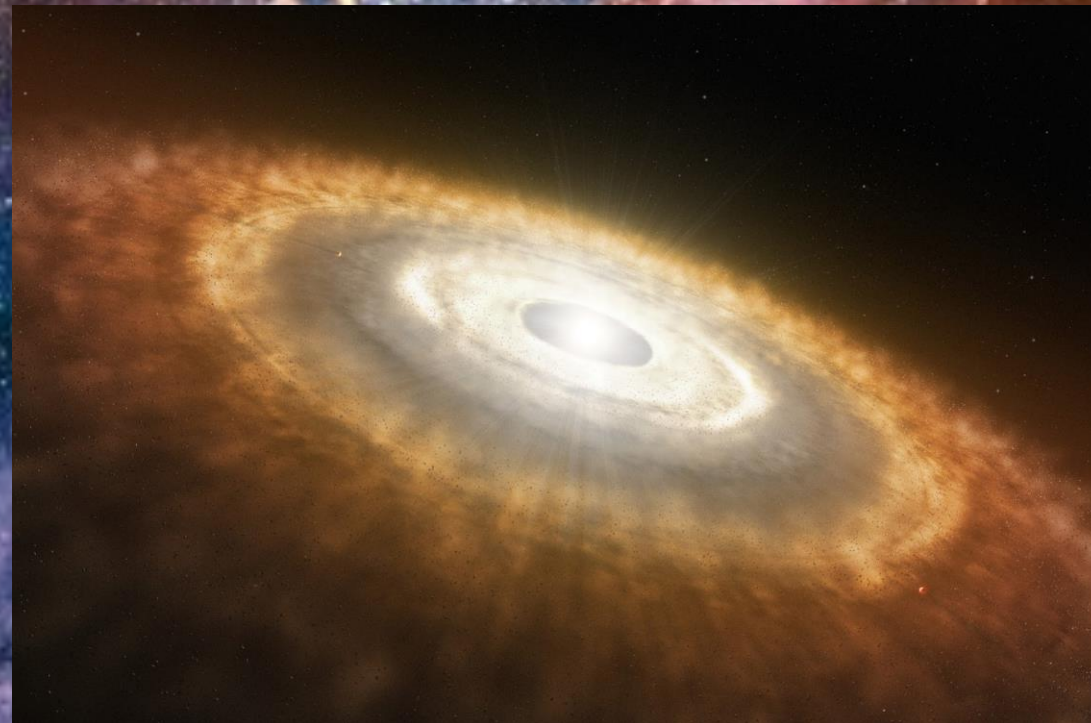
$$v = \sqrt{\frac{Gm}{r}}$$

v : velocity of the molecular

G : gravitational constant

m : protostellar mass

r : radius at which the velocity
is being measured



Keplerian curve

- ▶ A kind of curve that obeys Kepler's law, and is inversely proportional to the square root of its distance from the center
- ▶ It can be used to measure the mass of the source

column density

- ▶ It means the quantity of molecular or matter per unit area, and we can use it to times the area of the disk to estimate the mass of the disk
- ▶ Besides ,column density is based on radiative transfer and collisional excitation of molecules

column density

The column density N is equal to

$$N_{tot}(\text{C18O}) = \frac{3 \cdot h}{8 \cdot \pi^3 \cdot u^2 \cdot J_u \cdot R_i} \left(\frac{k \cdot T_{ex}}{h \cdot B} + \frac{1}{3} \right) \exp\left(\frac{E_u}{k \cdot T_{ex}}\right) \left[\exp\left(\frac{h \cdot \nu}{k \cdot T_{ex}}\right) - 1 \right]^{-1} \int dv \tau_\nu$$

$$E_u = 15.81 \text{K}(\text{C18O}, 2-1)$$

$$T_{ex} = 10 \text{K}$$

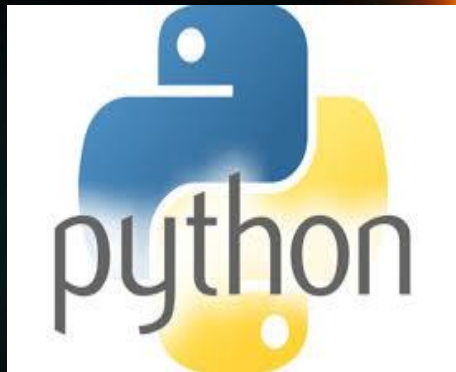
$$u = 1.1079 \cdot 10^{-19}$$

$$J_u = 2(\text{C18O}, 2-1)$$

$$R_i = 1$$

$$B = 54891.42 \text{(MHz)}$$

program



Read header

Calculate column
density

Draw keplerian
curve

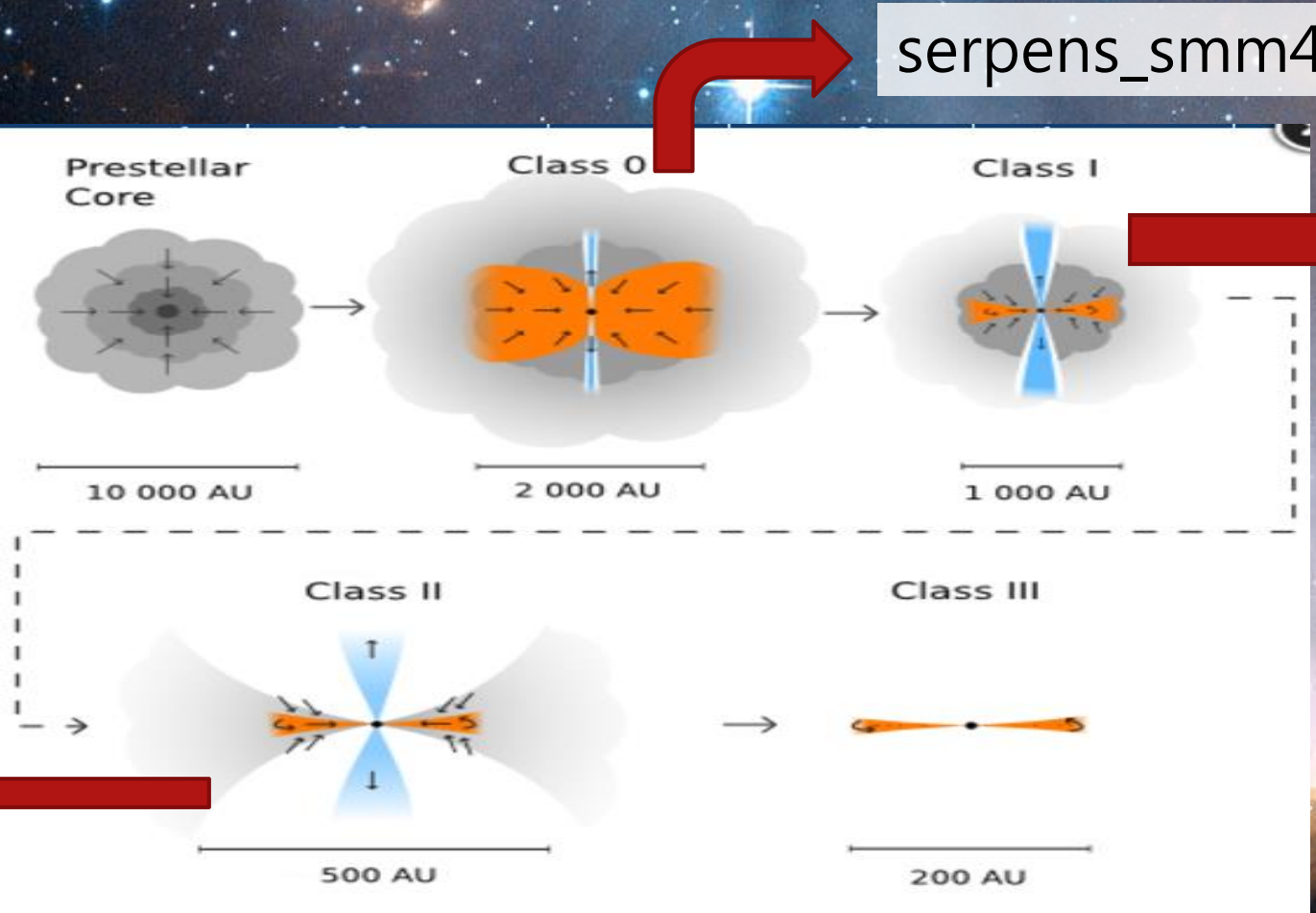
casa

Moment map

Position-velocity
diagram

Importfits and
exportfits

Star formation



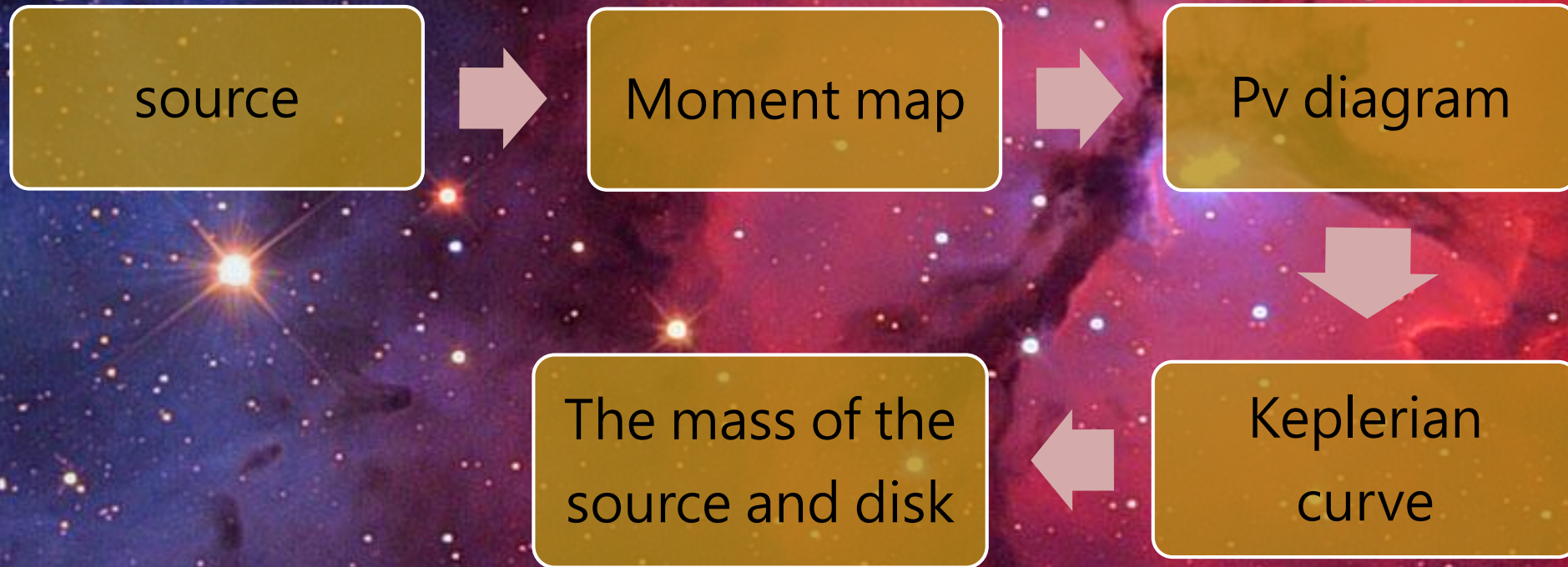
serpens_smm4

DG tau B

IRS_48



outline2



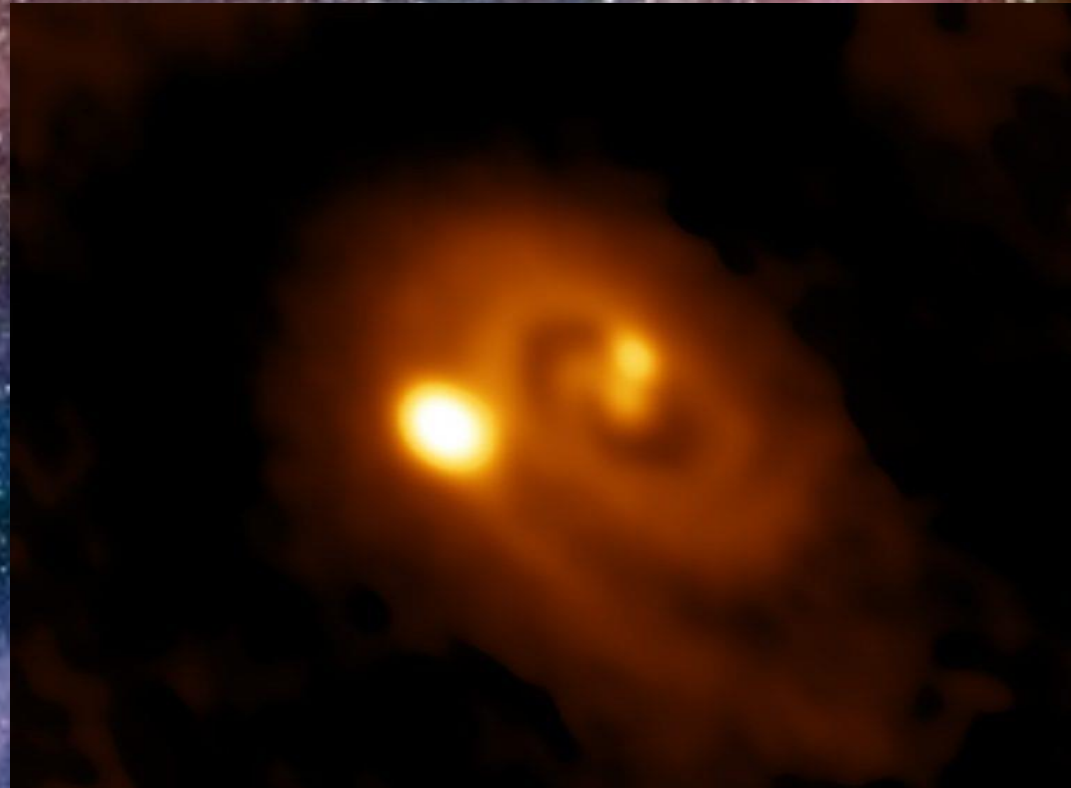
Example: L1448

center:03h25m38.8
30d44m04

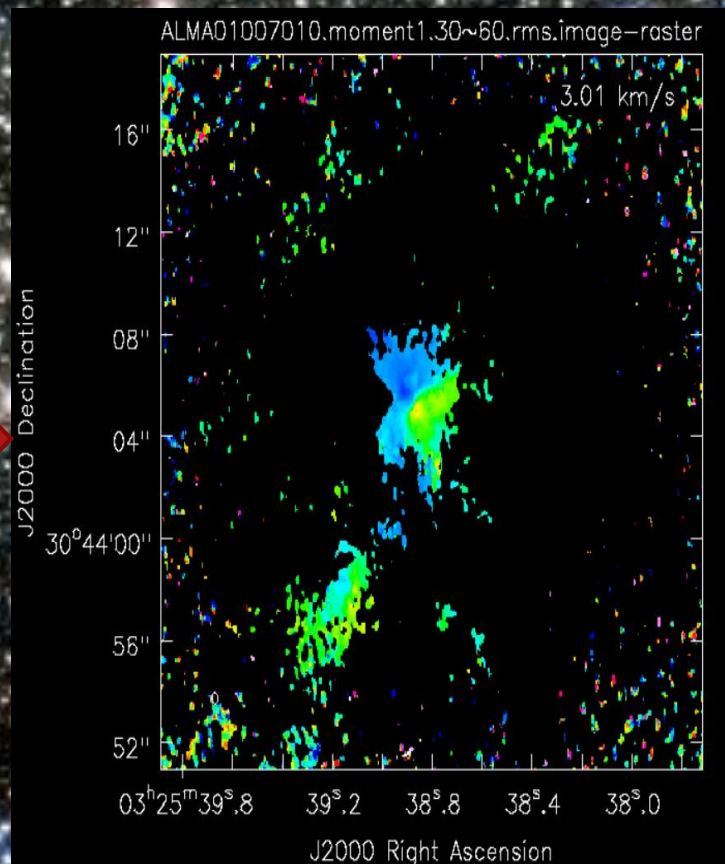
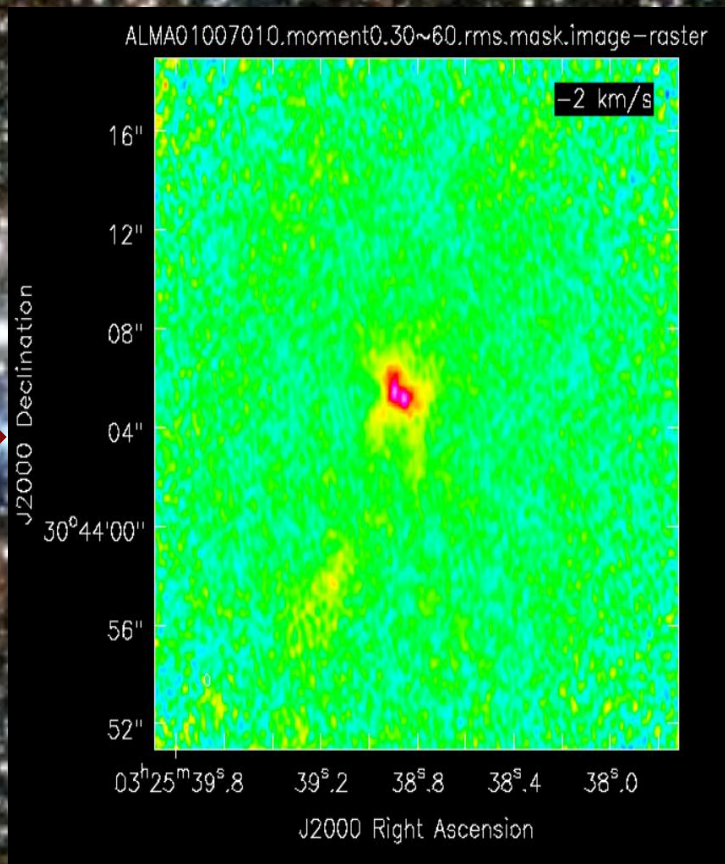
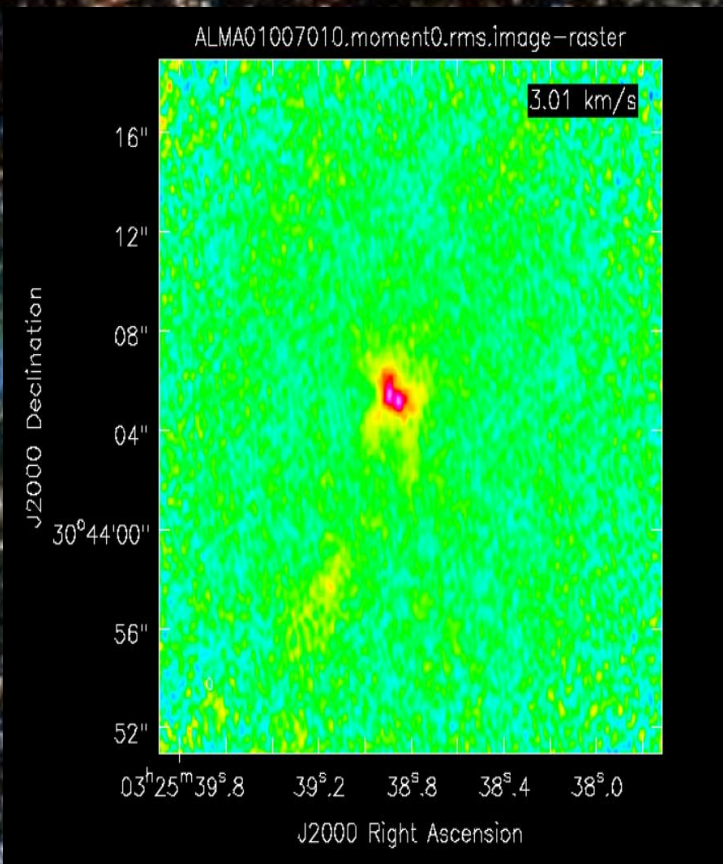
Class 0

System: triple protostar
System

Spectrum: C18O(2-1)

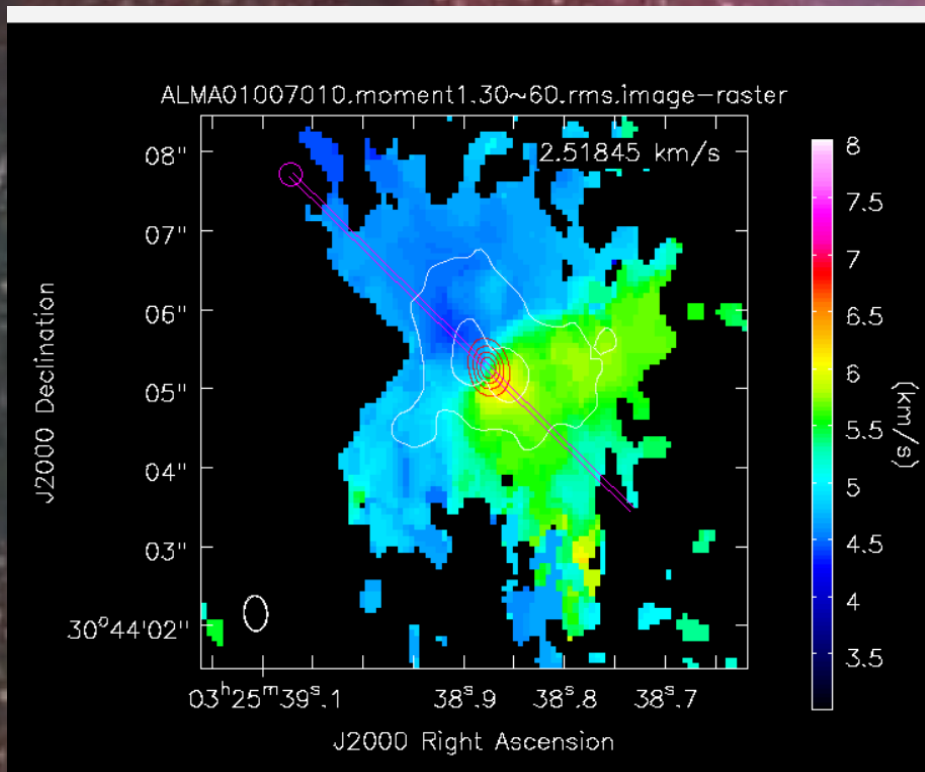


Moment map

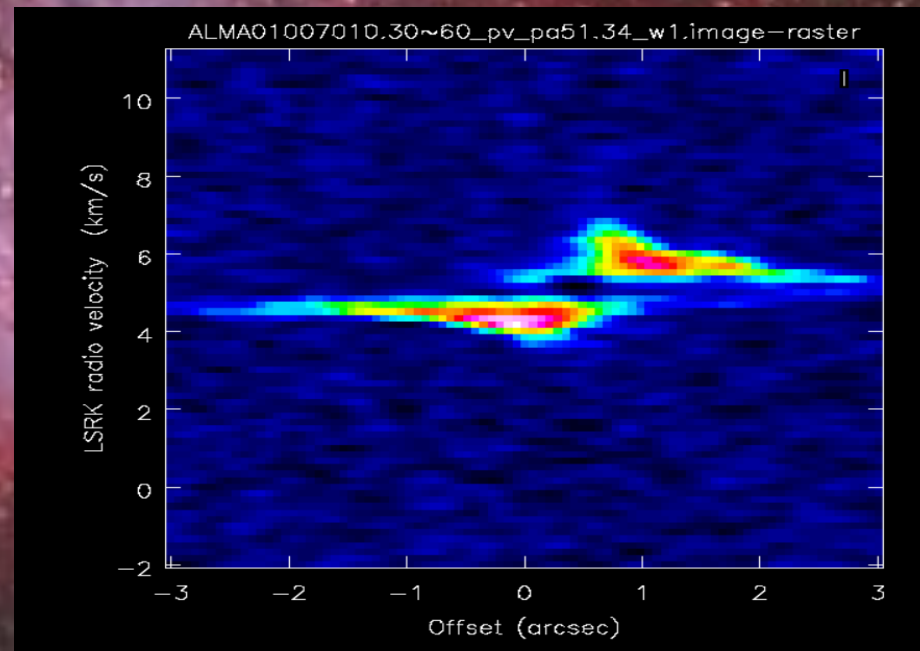


Position-velocity diagram

(1) We can use it to diagnose whether it is a keplerian disk

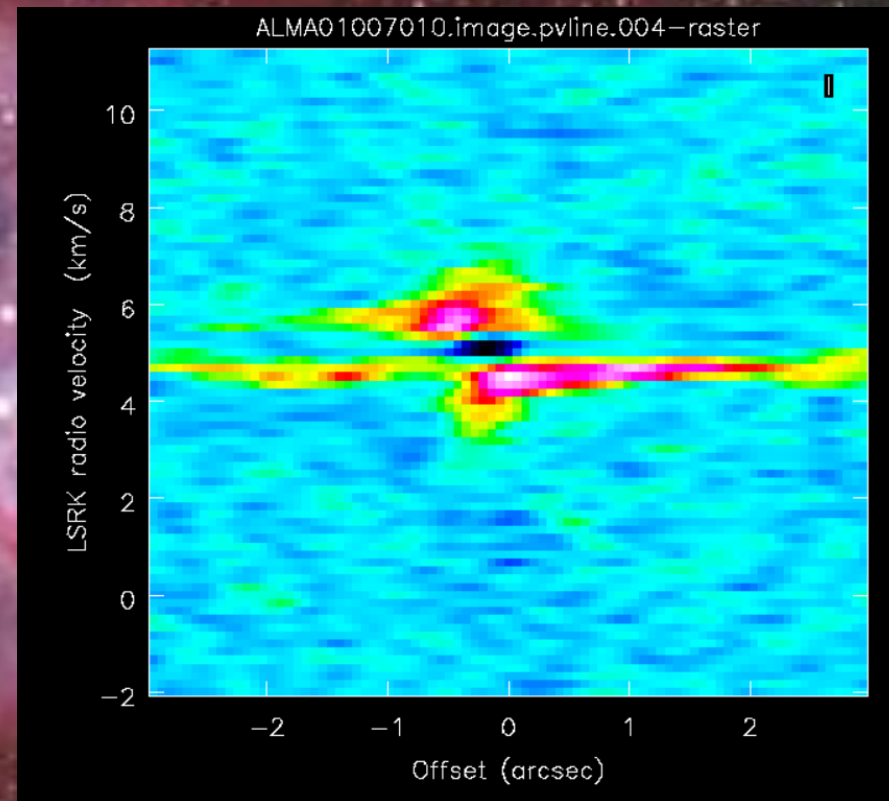
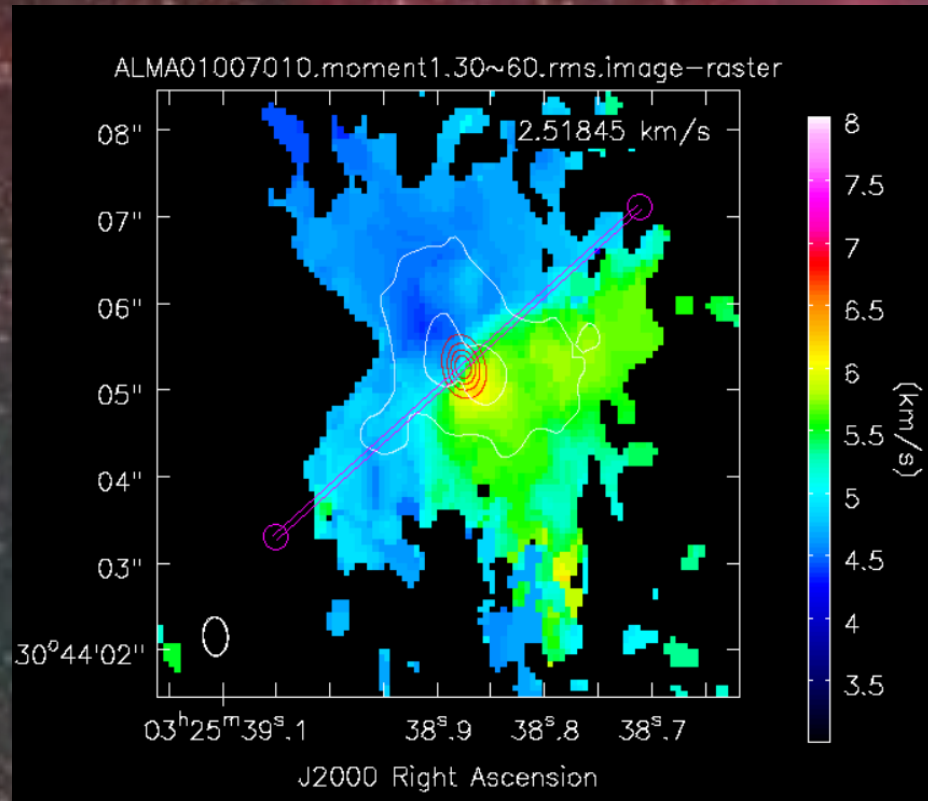


Pa=51.34

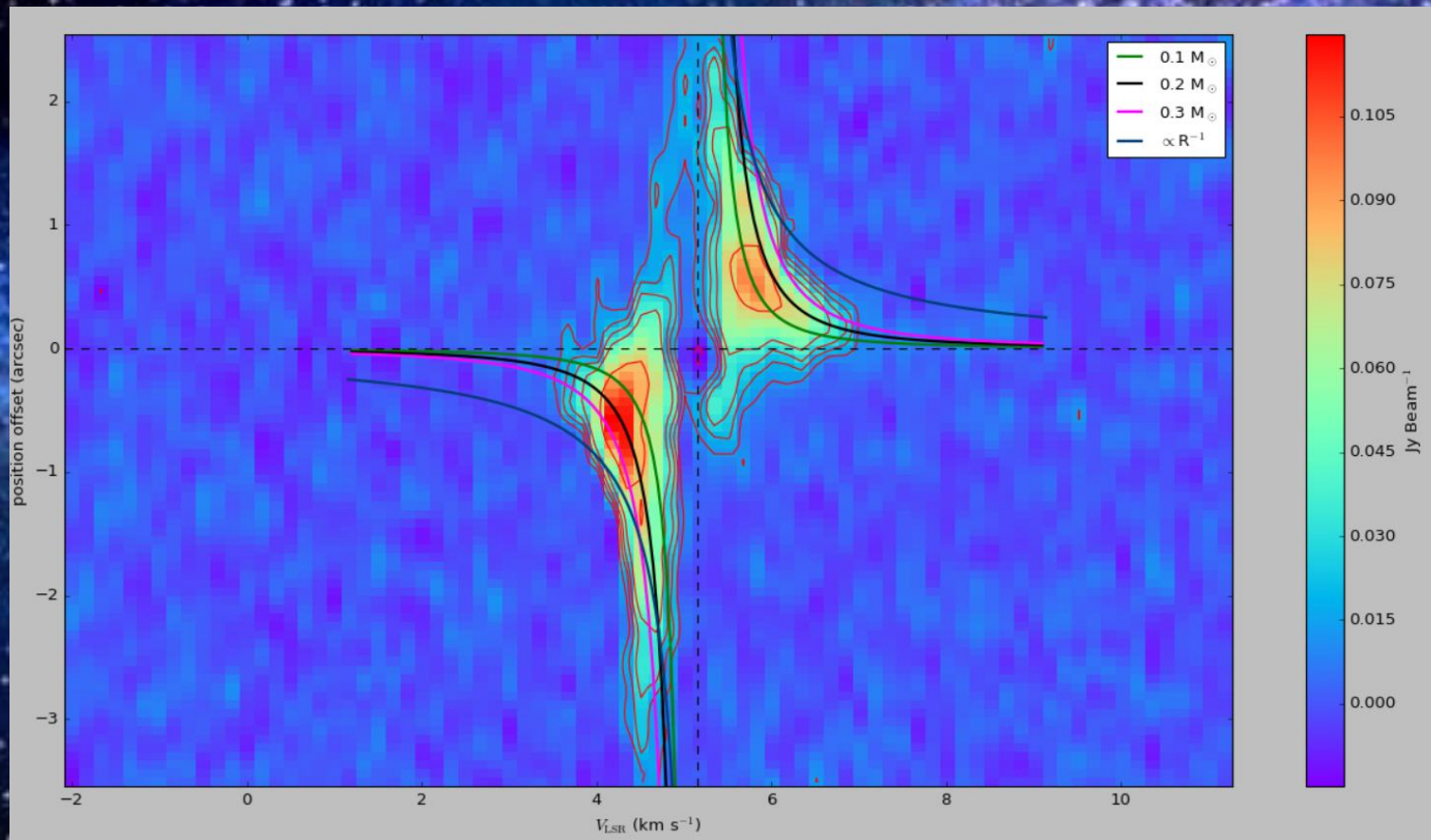


Position-velocity diagram

$P_a = -40.91$



Keplerian curve



The mass of the disk

- ▶ I use column density to calculate the quantity of the molecular per cm^2 , and use this to calculate the mass of the disk.
- ▶ Therefore I calculate the mass of the disk(L1448) for 0.045 solar mass.

Example:Lupus3 MMS

System: fourth
protostar system

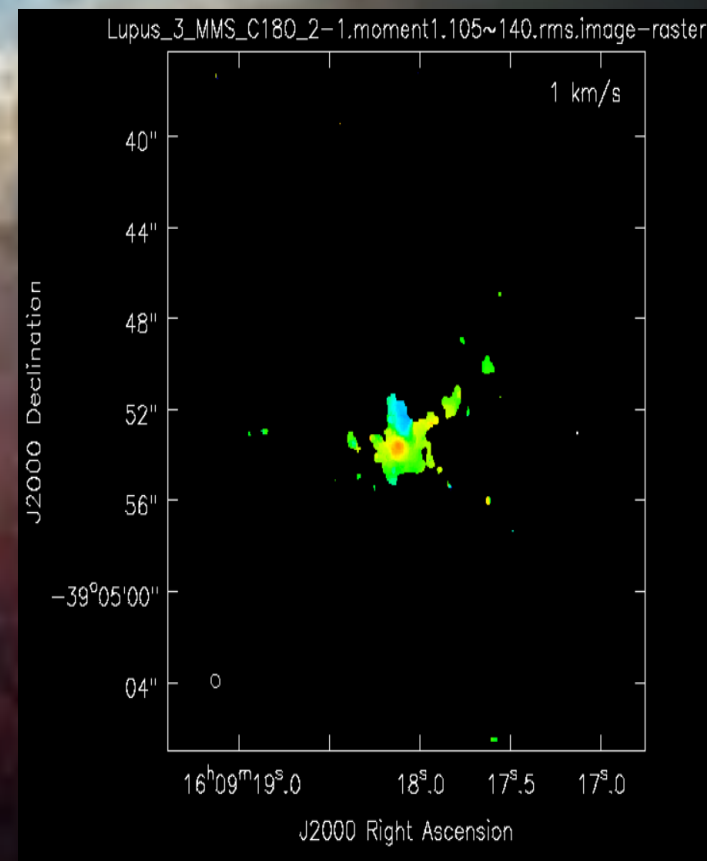
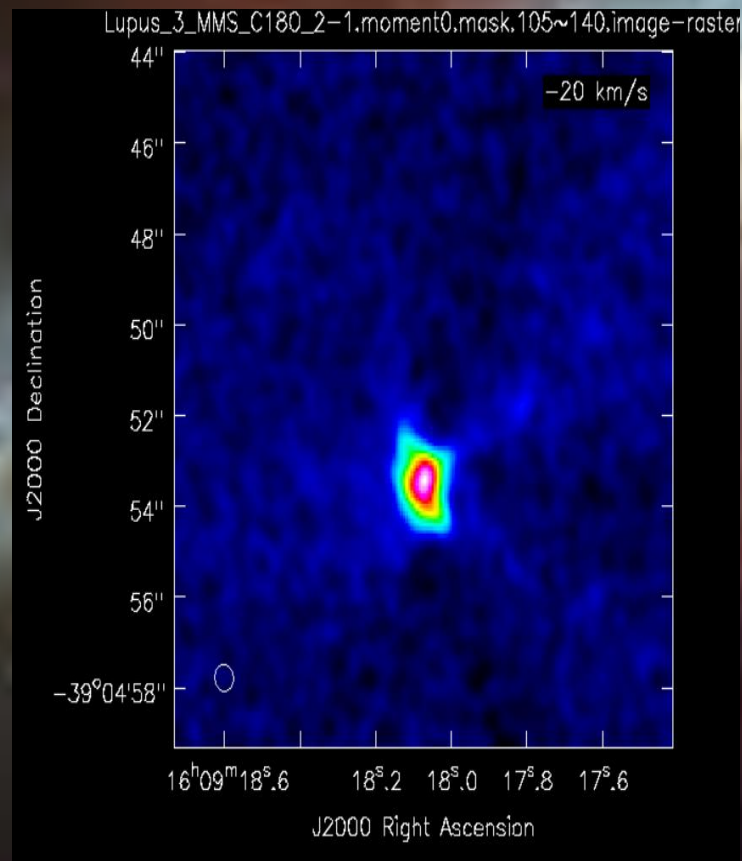
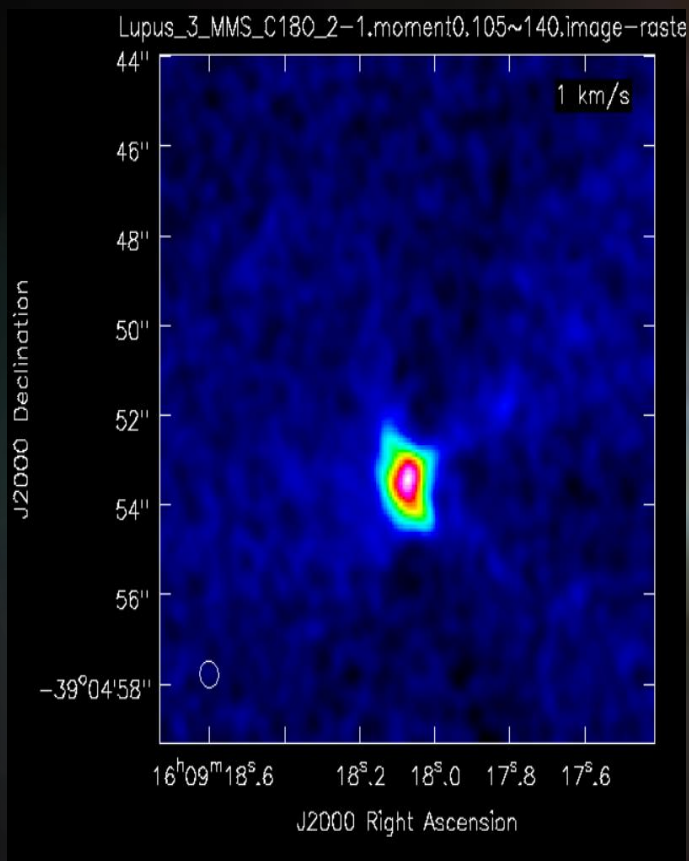
Center:4h31m44.5
18d8

Class:0

Spectrum:C18O(2-1)

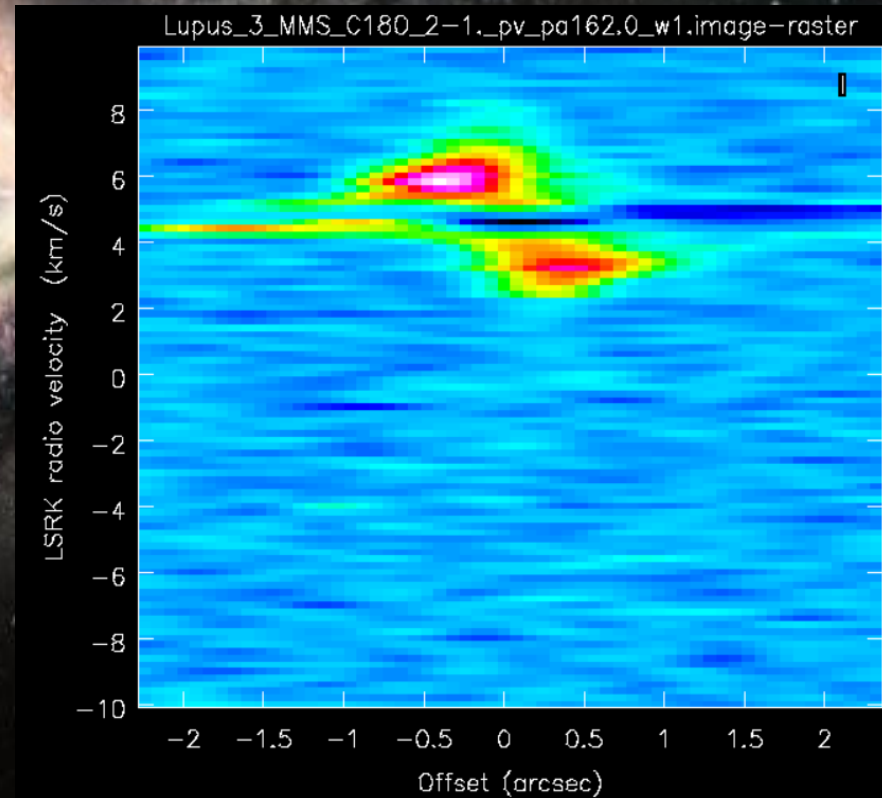
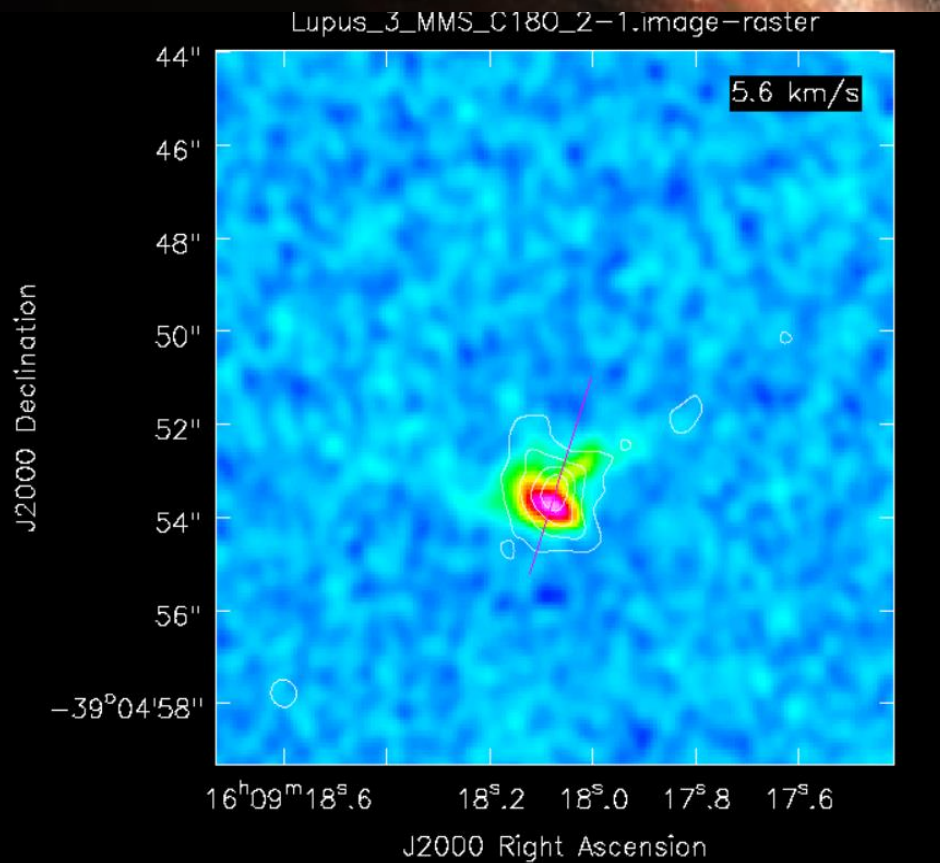


Moment map



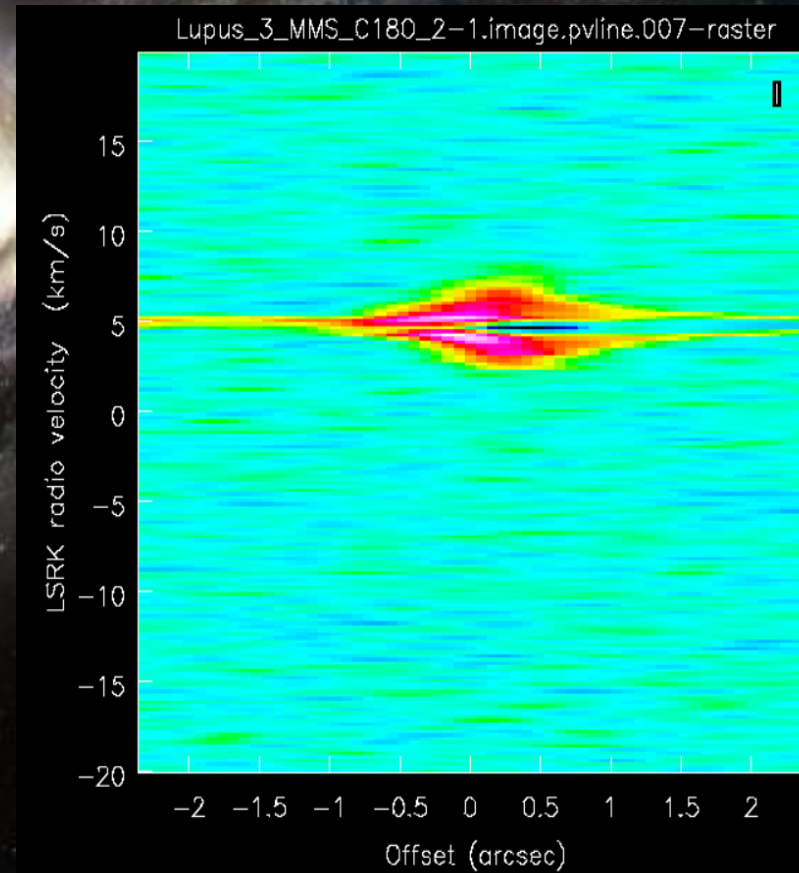
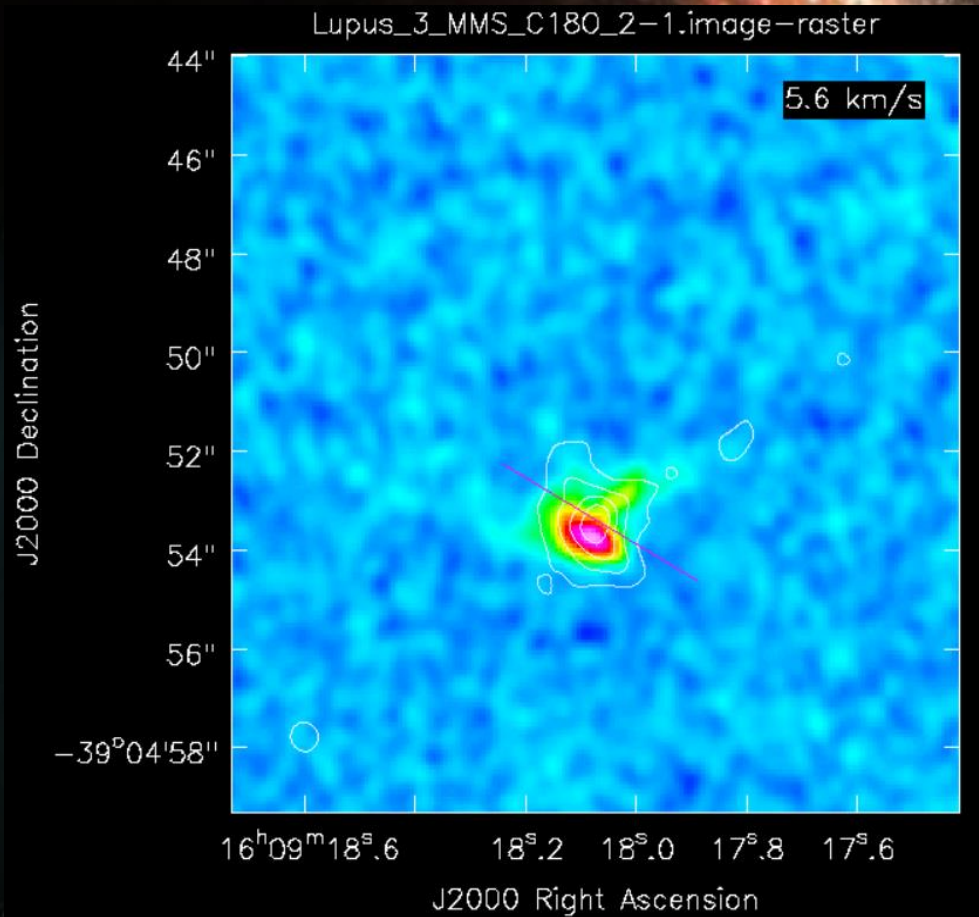
Pv diagram

Pa=162

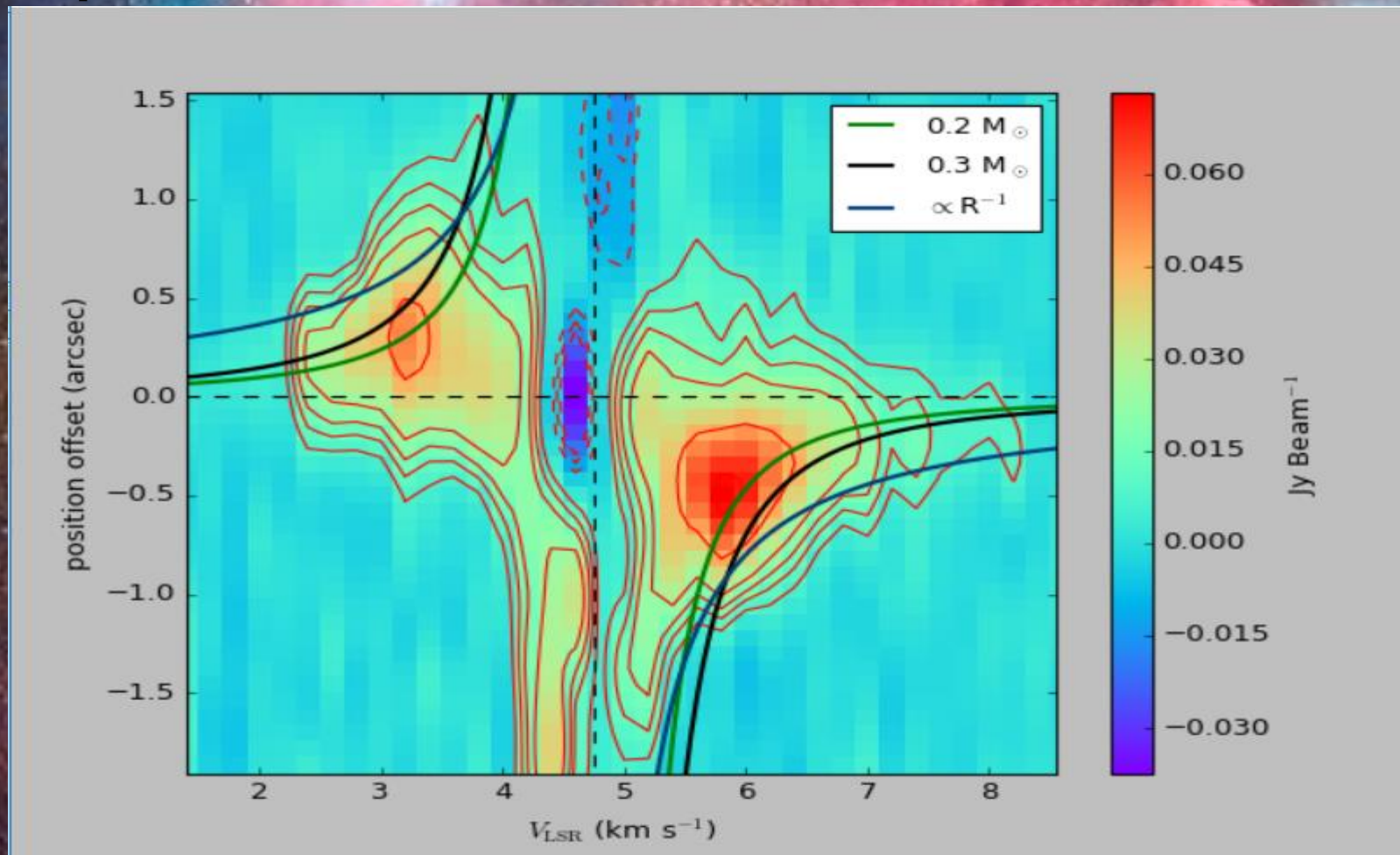


Pv diagram

Pa=60



Keplerian curve



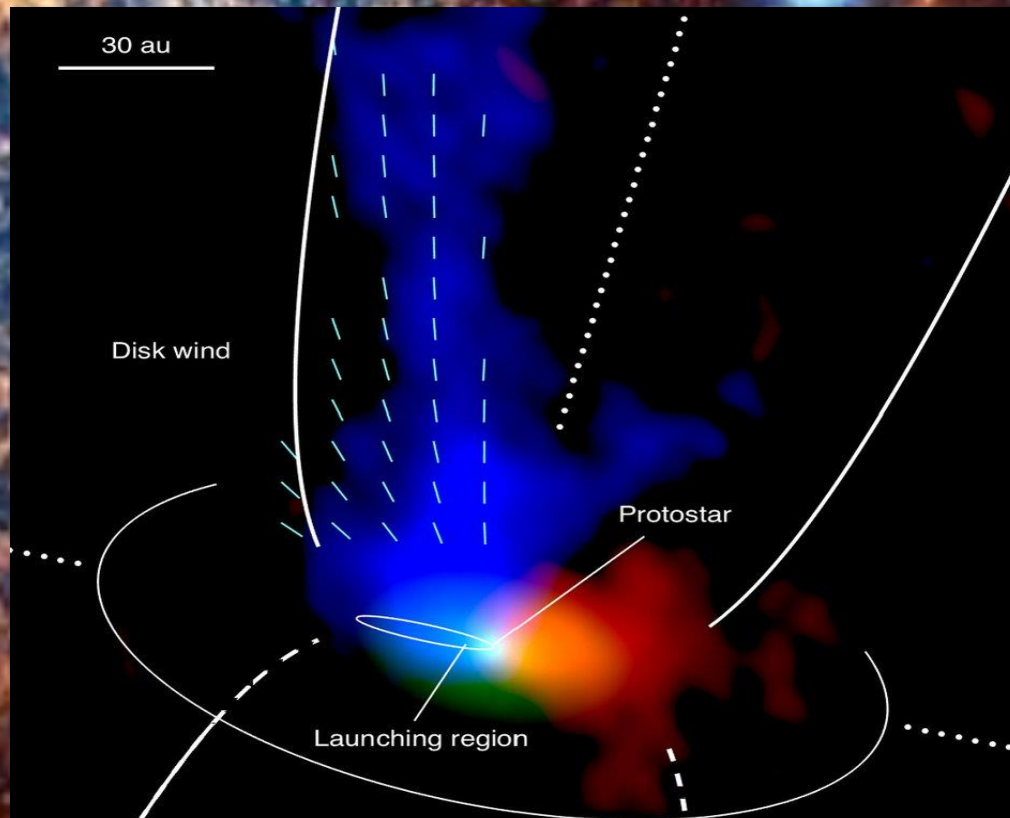
Example:TMC1A

System: binary
protostar system

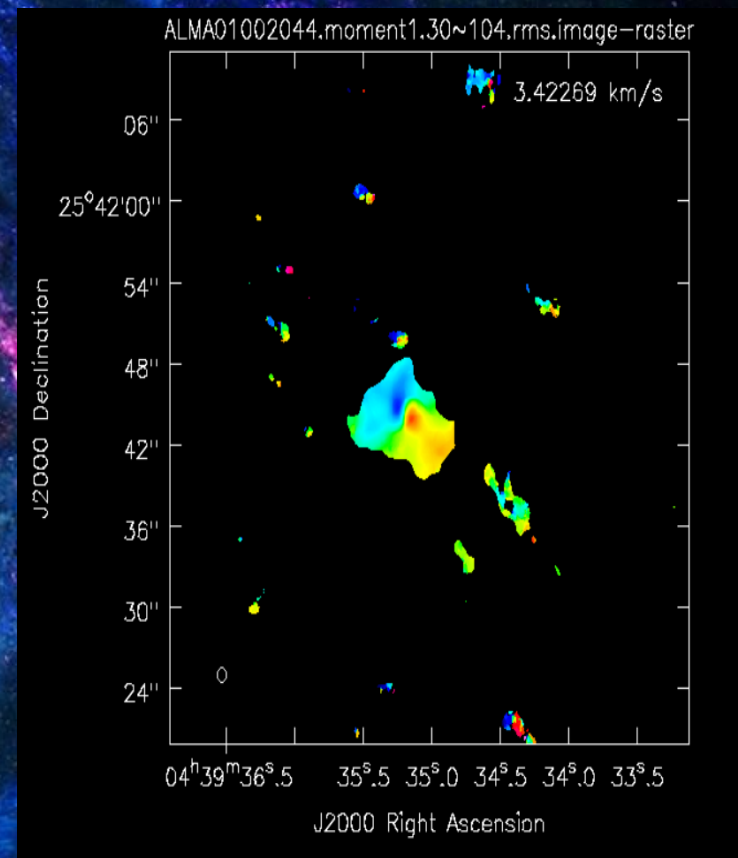
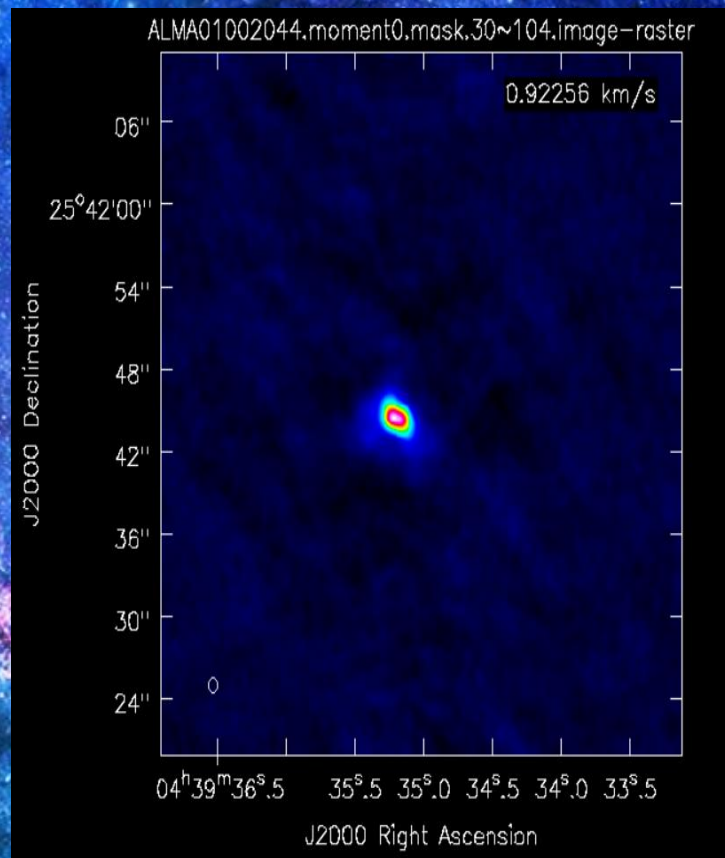
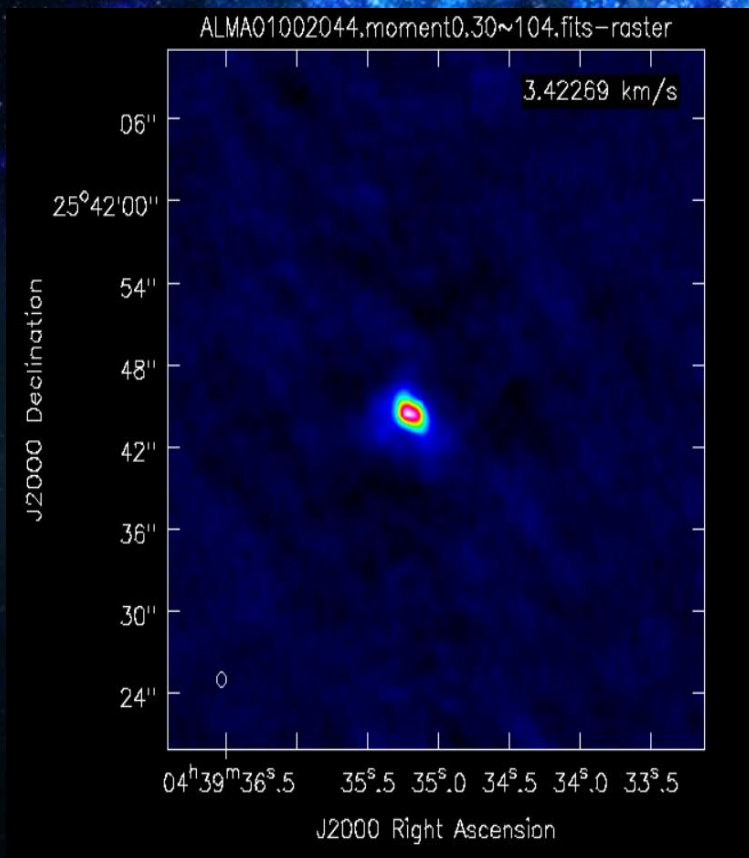
Center:04h39m53.9
+26d03m09

Class:1

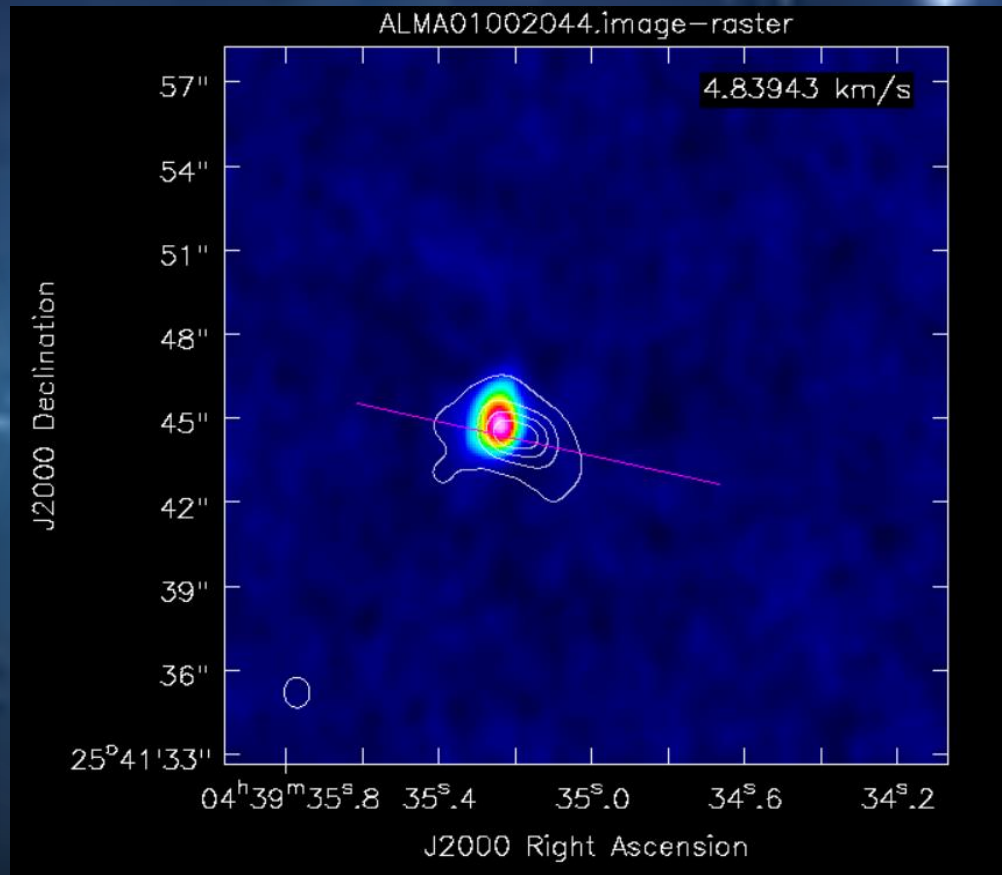
Spectrum:C18O(2-1)



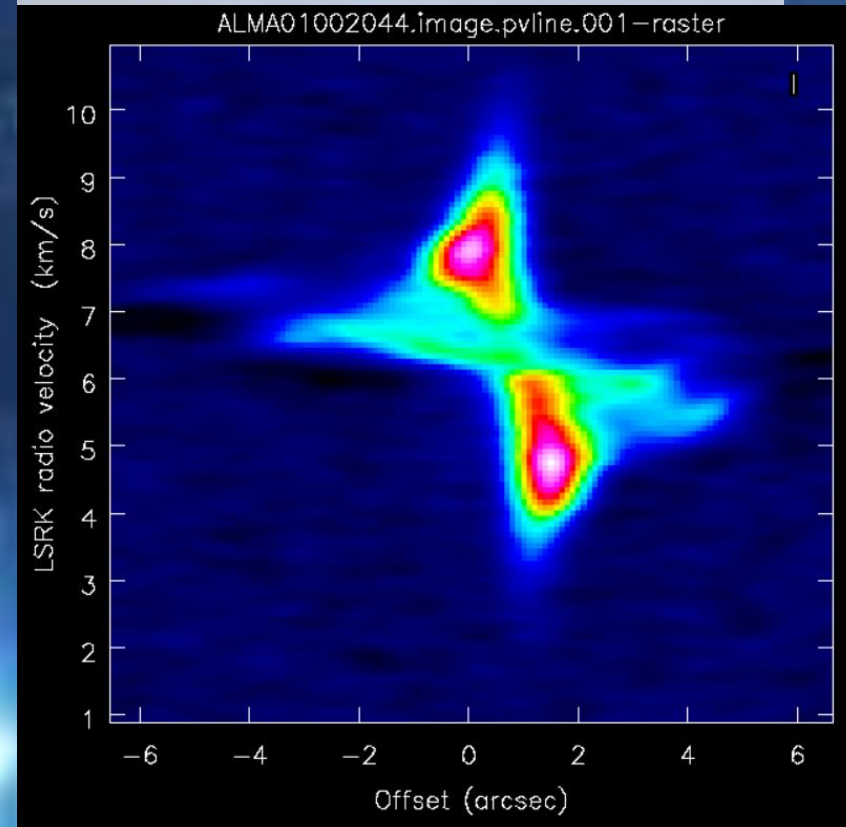
Moment map



Pv diagram

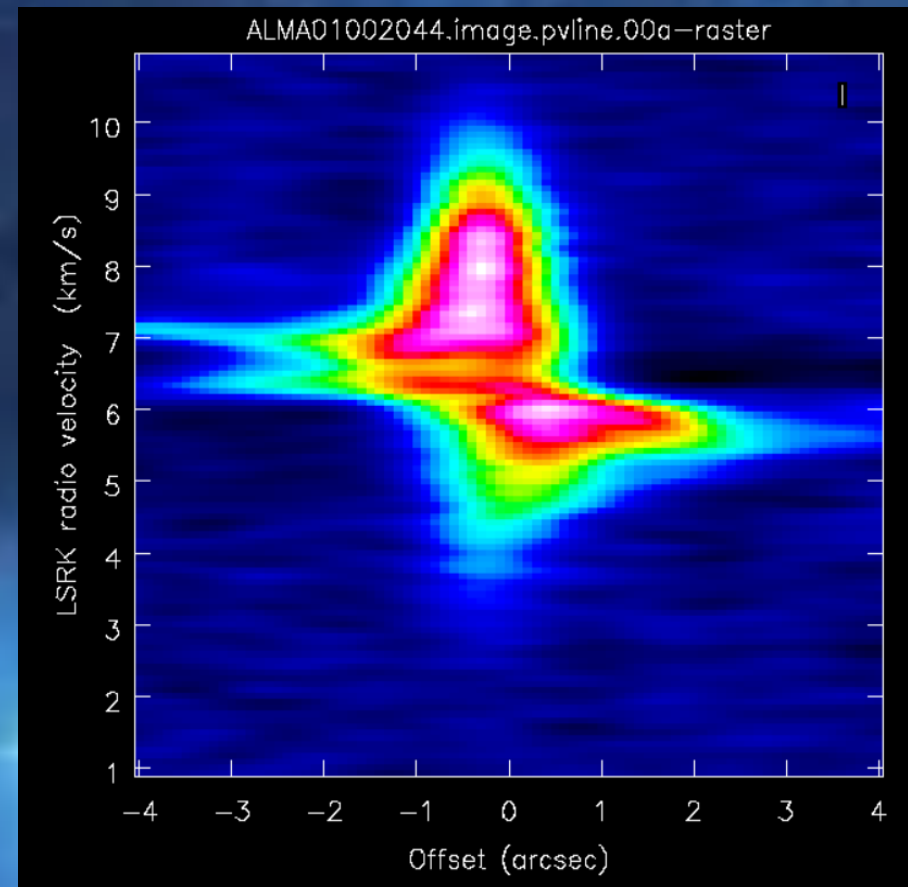
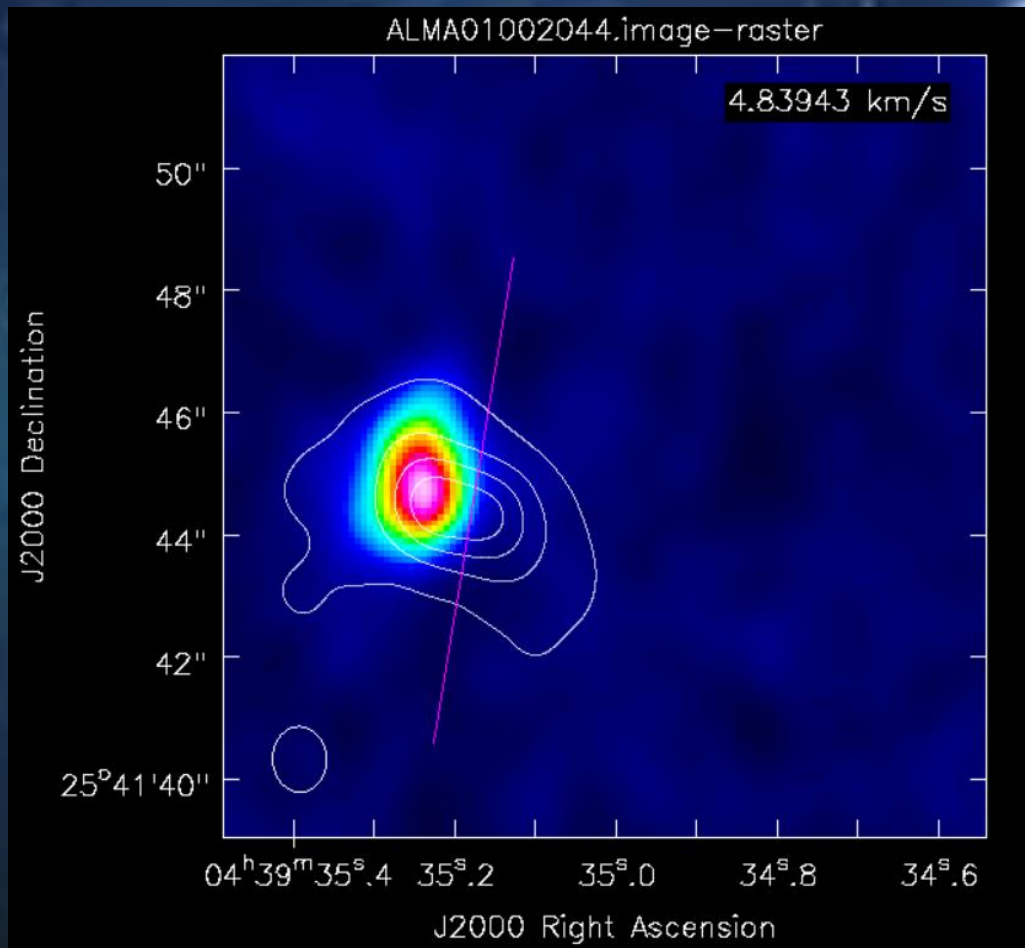


Pa=77.3

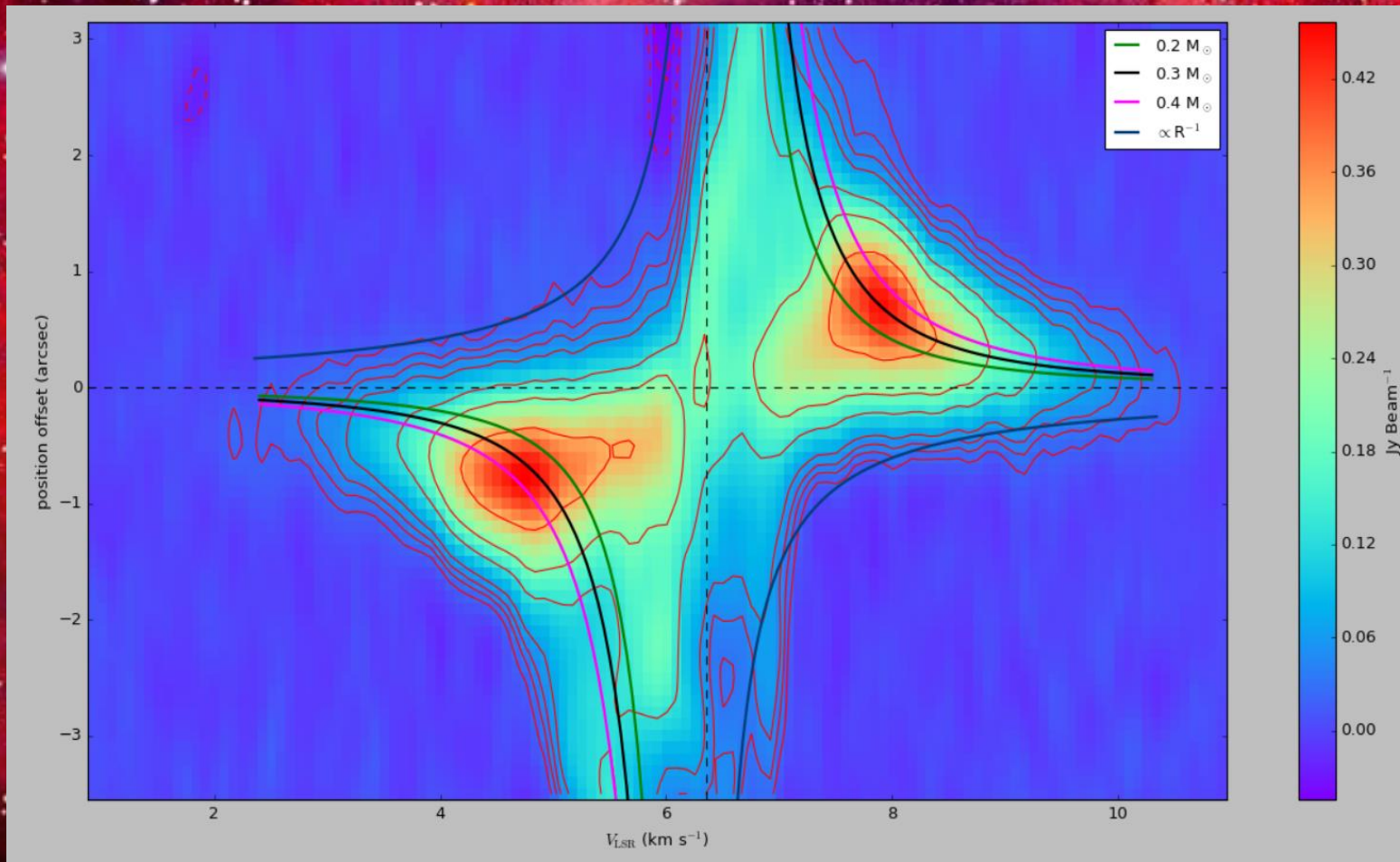


Pv diagram

Pa=170.4



Keplerian curve



The mass of the disk

- ▶ I use column density to calculate the quantity of the molecular per cm^2 , and use this to calculate the mass of the disk.
- ▶ Therefore I calculate the mass of the disk (TMC1A) for 0.01875 solar mass.

Example: DG TAU B

**System: single
protostar system**

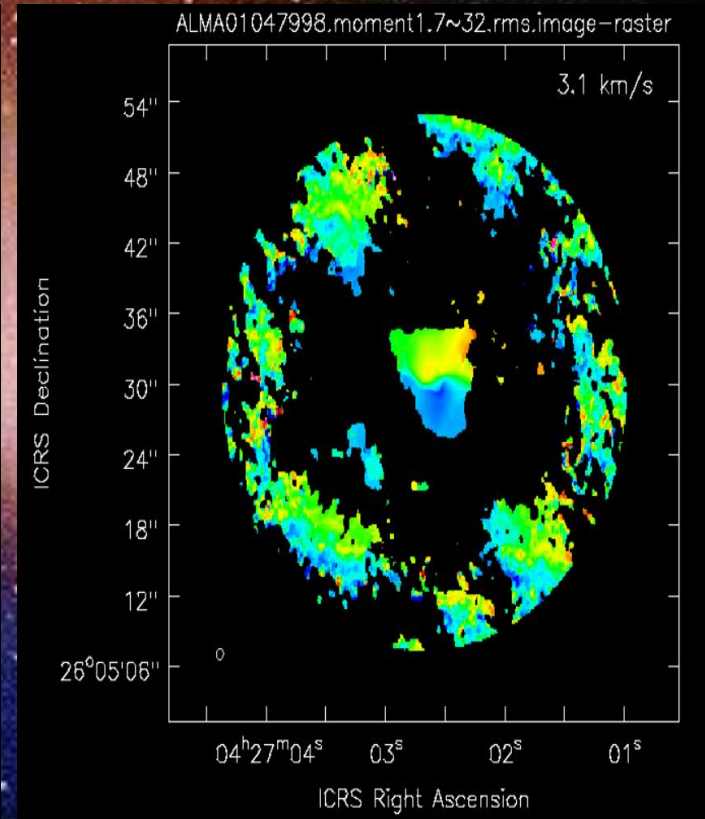
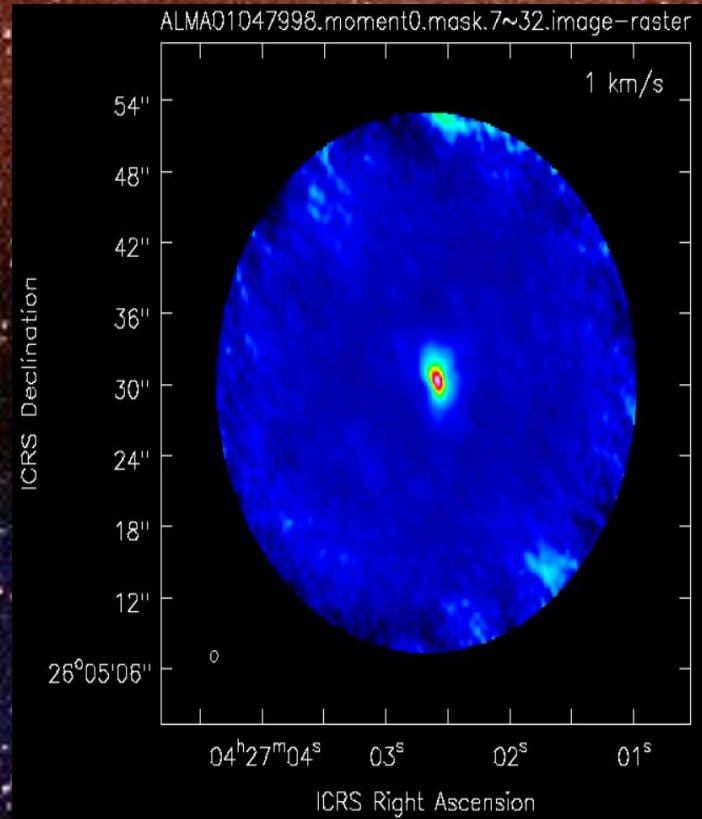
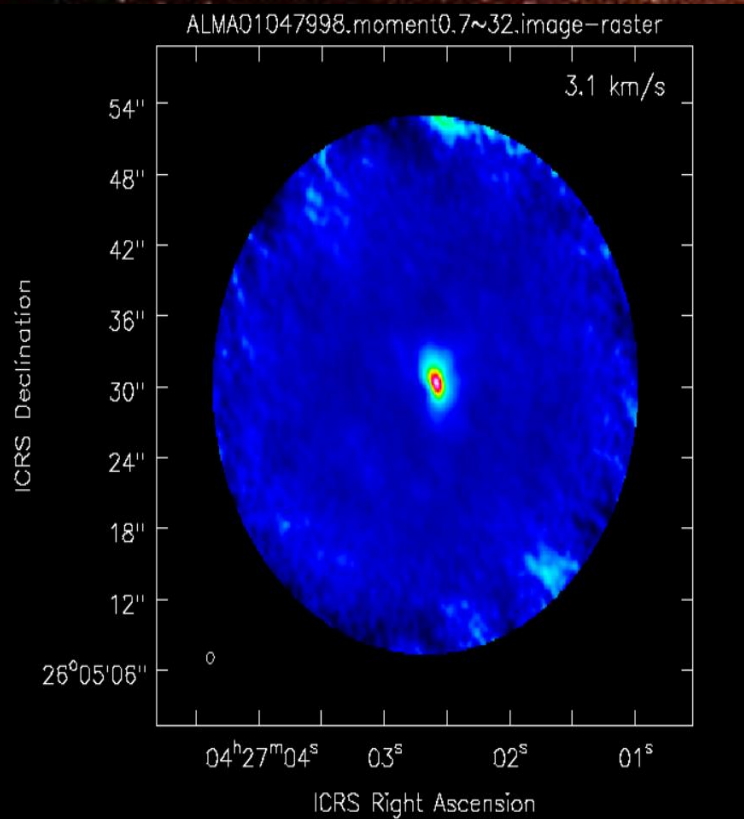
Center:04h27m02.6
+26d05m29

Class:1

Spectrum:13CO(2-1)

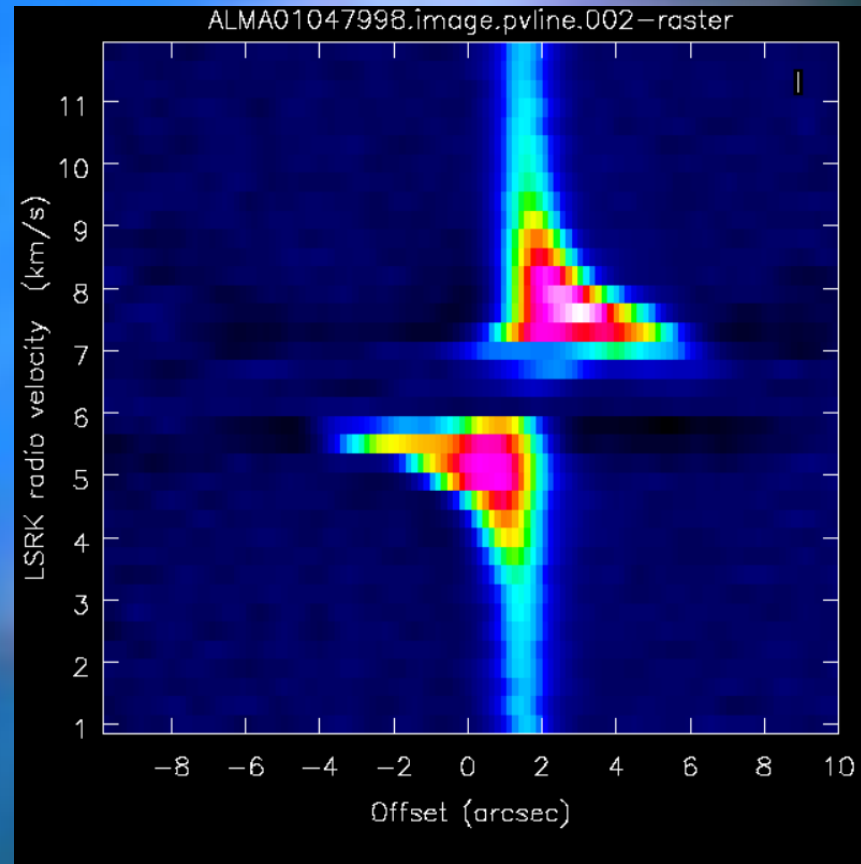
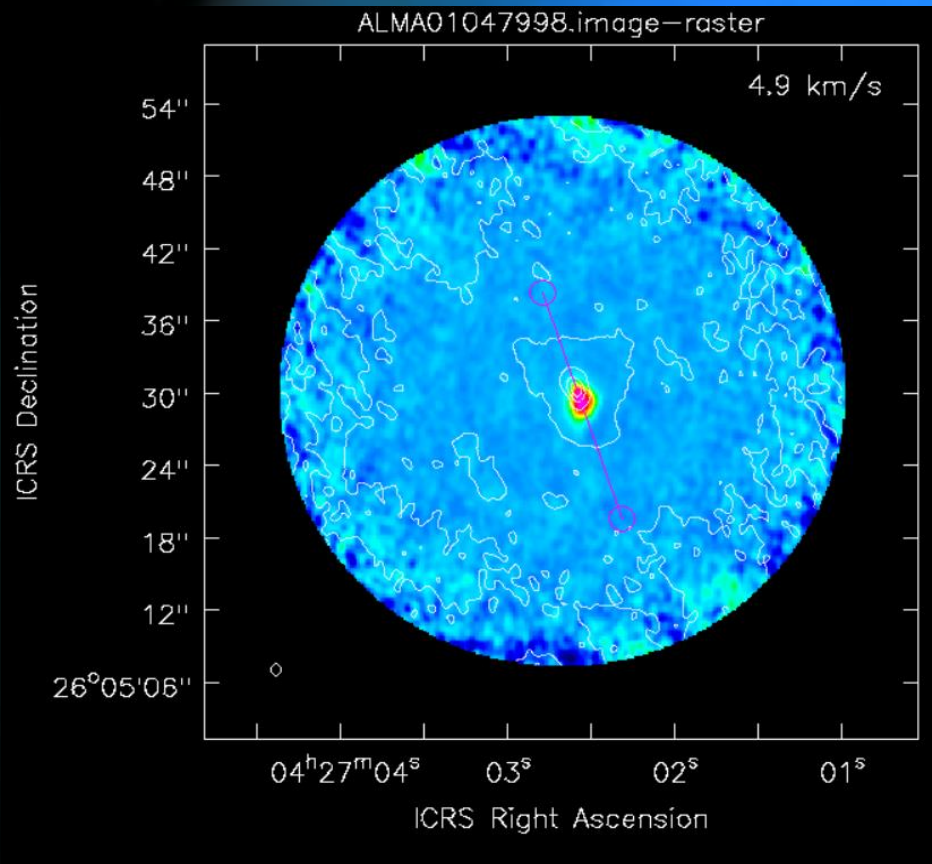


Moment map



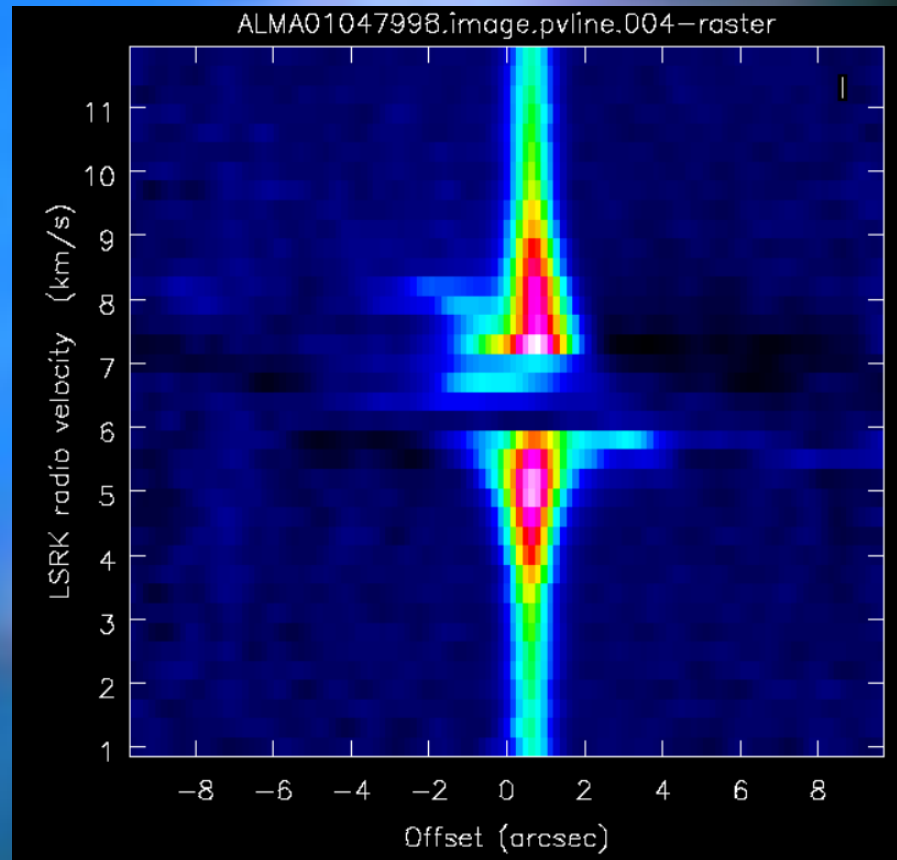
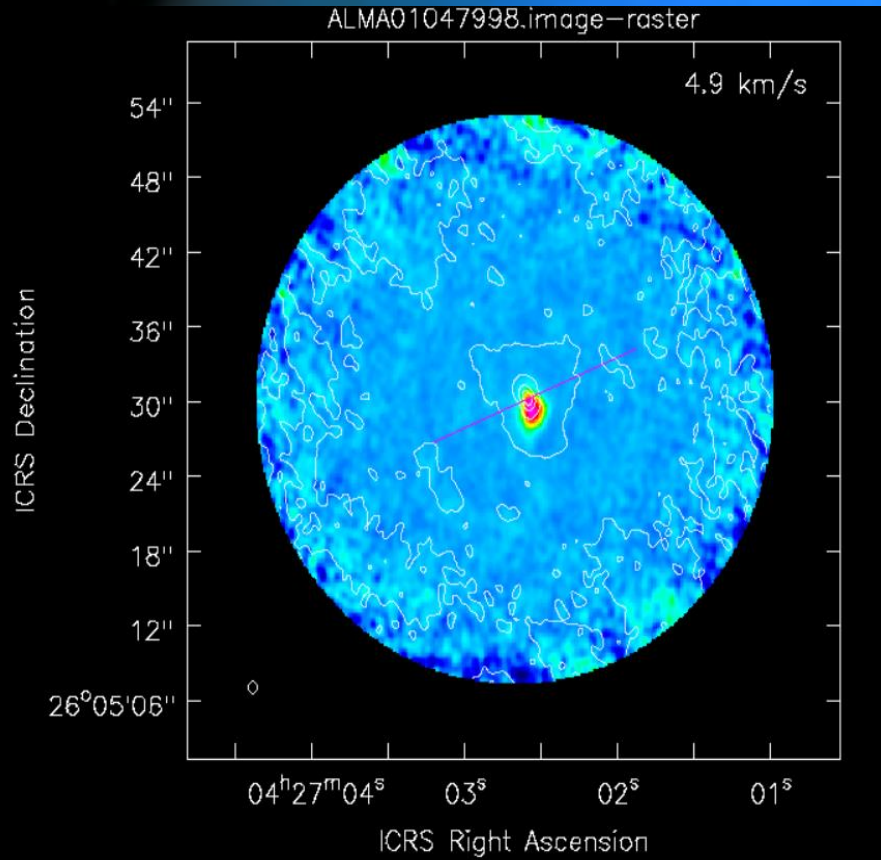
Pv diagram

Pa=18.64

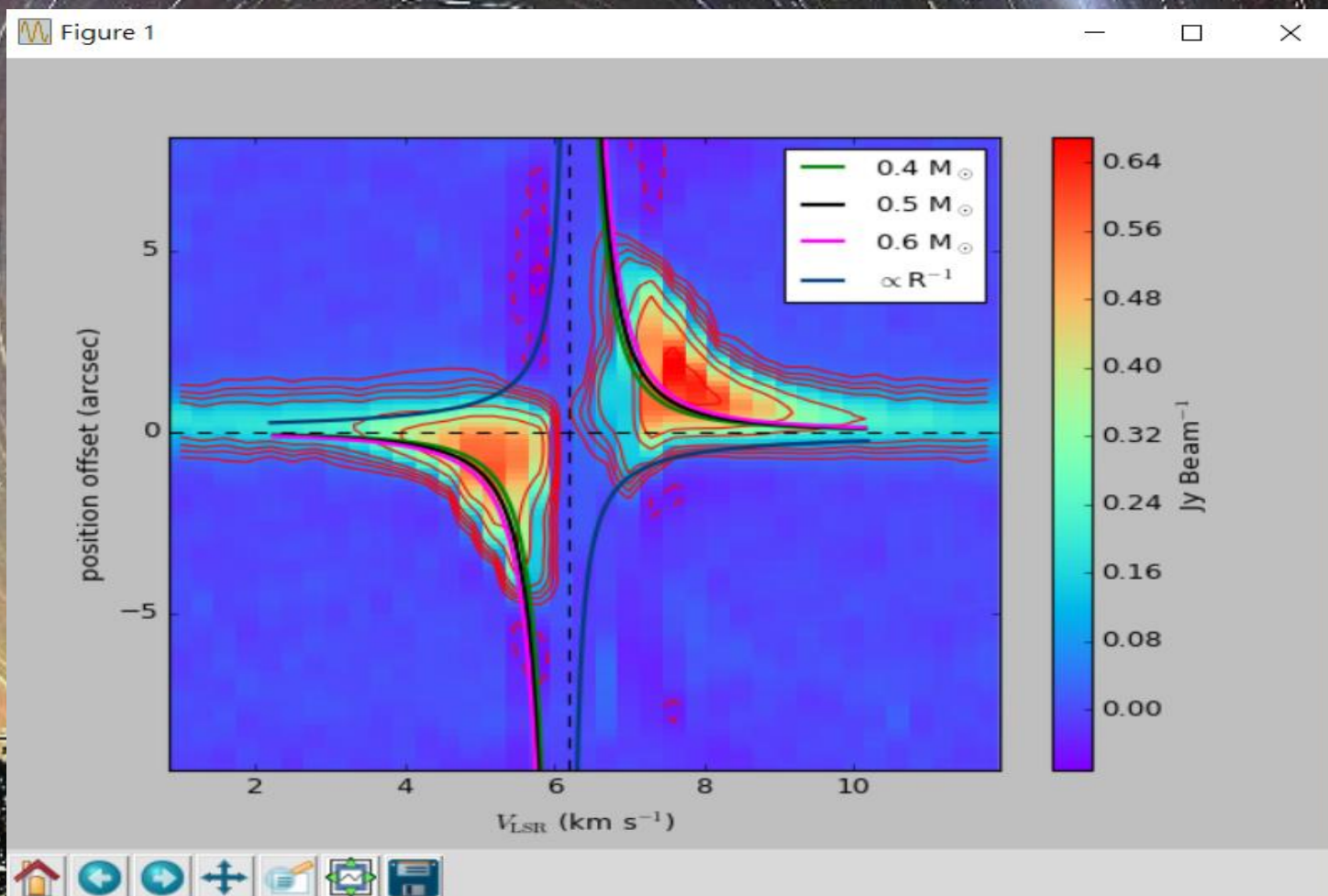


Pv diagram

Pa=-67



Keplerian curve



The mass of the disk

- ▶ I use column density to calculate the quantity of the molecular per cm^2 , and use this to calculate the mass of the disk.
- ▶ Therefore I calculate the mass of the disk (DG TAU B) for 0.0541 solar mass.

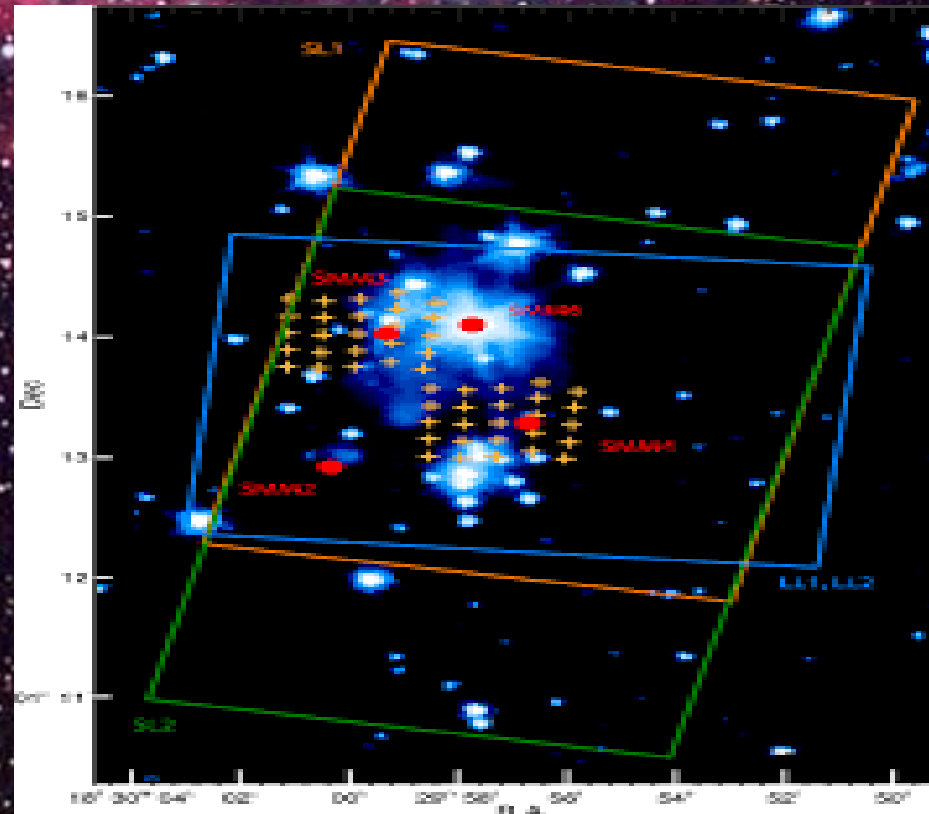
Example:serpens_smm4

System: single
protostar system

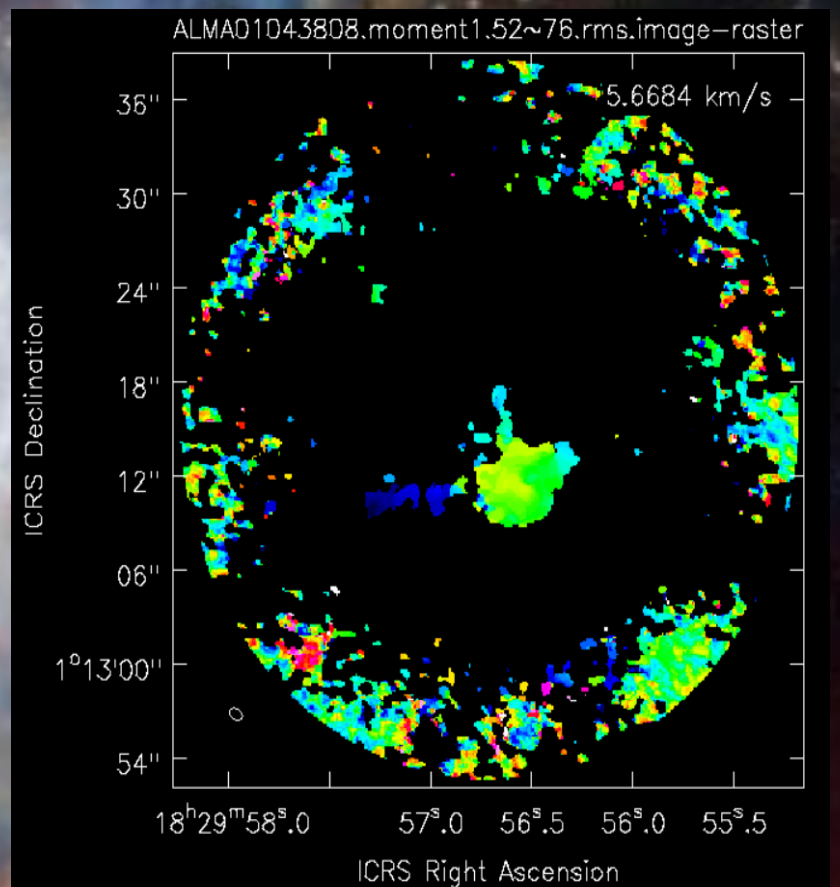
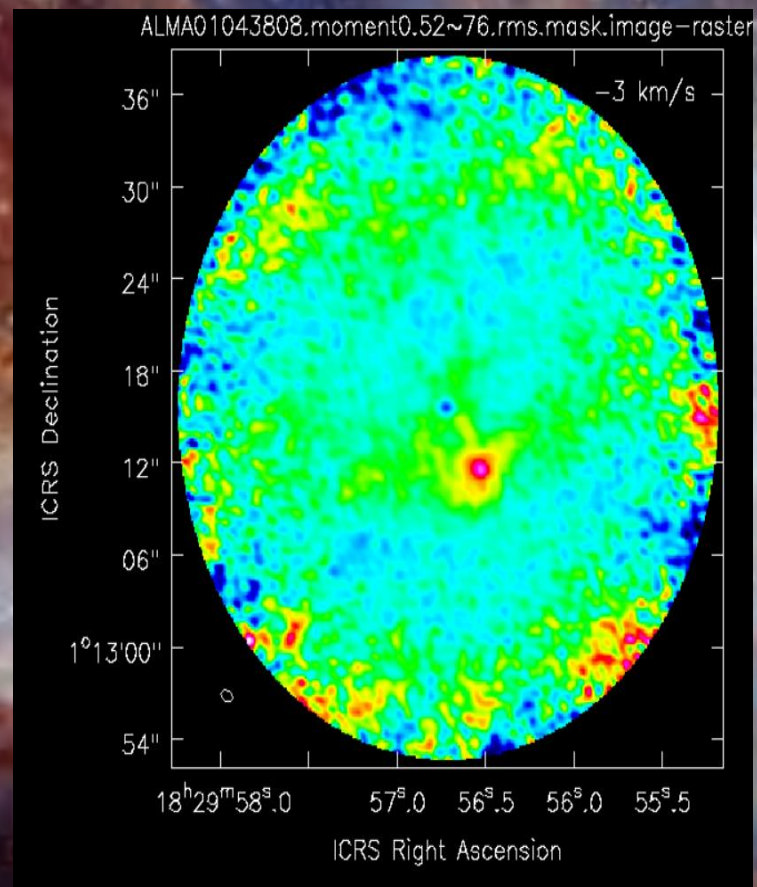
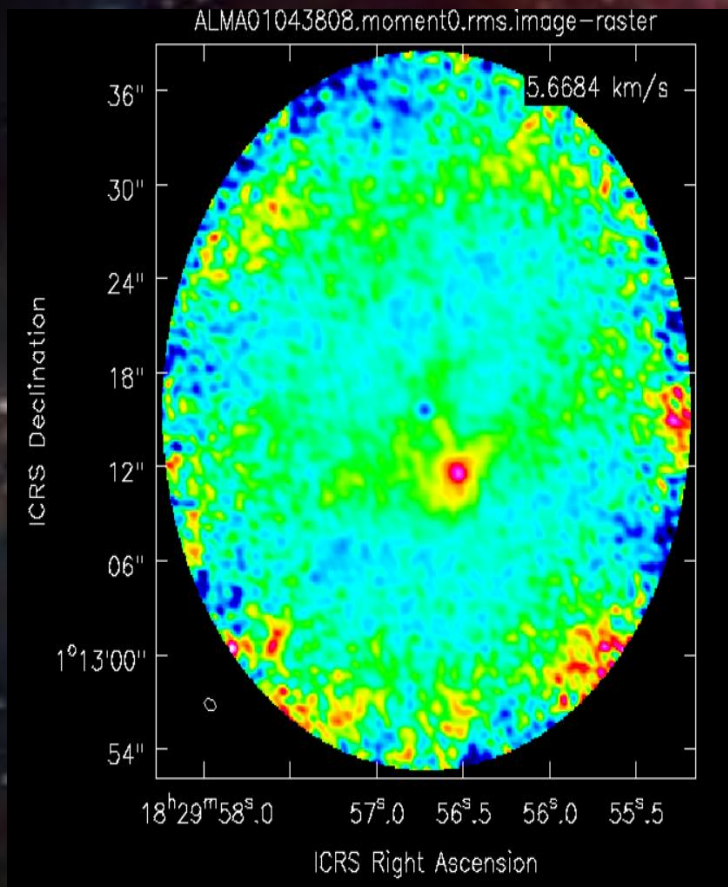
Center:18h29m56.7
+01d13m15

Class:0

Spectrum:C18O(2-1)

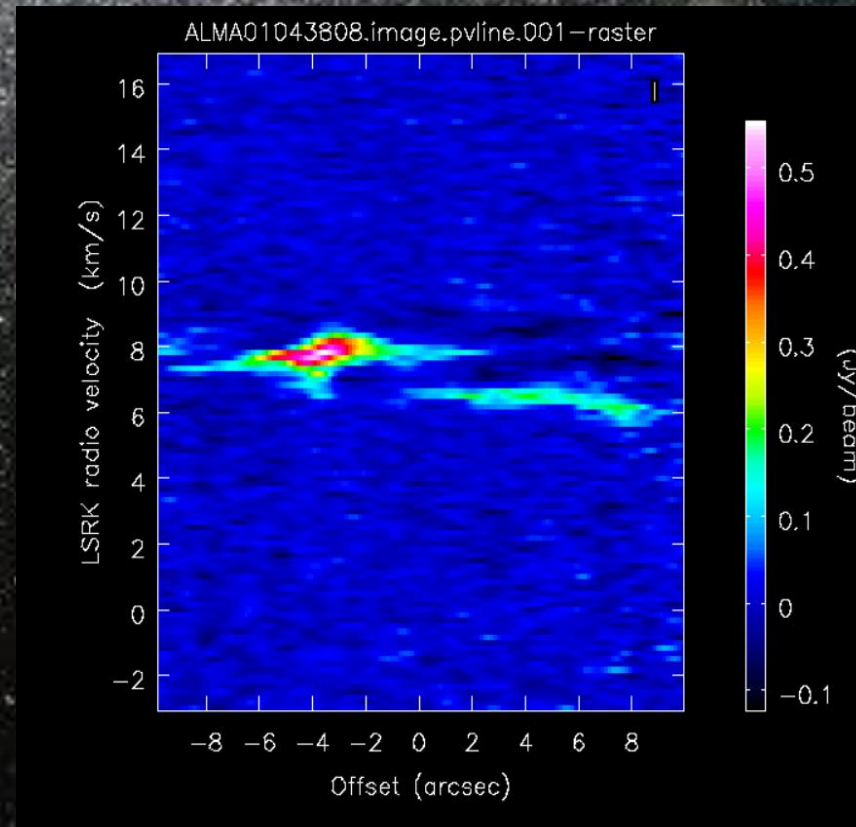
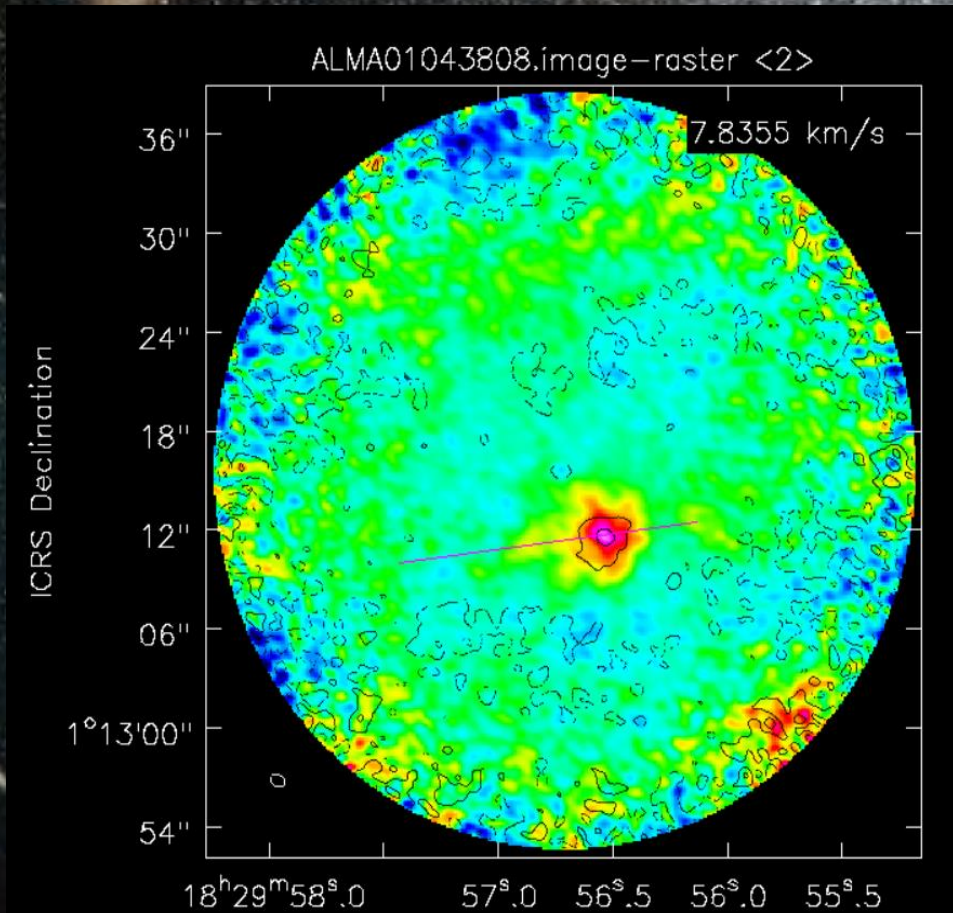


Moment map

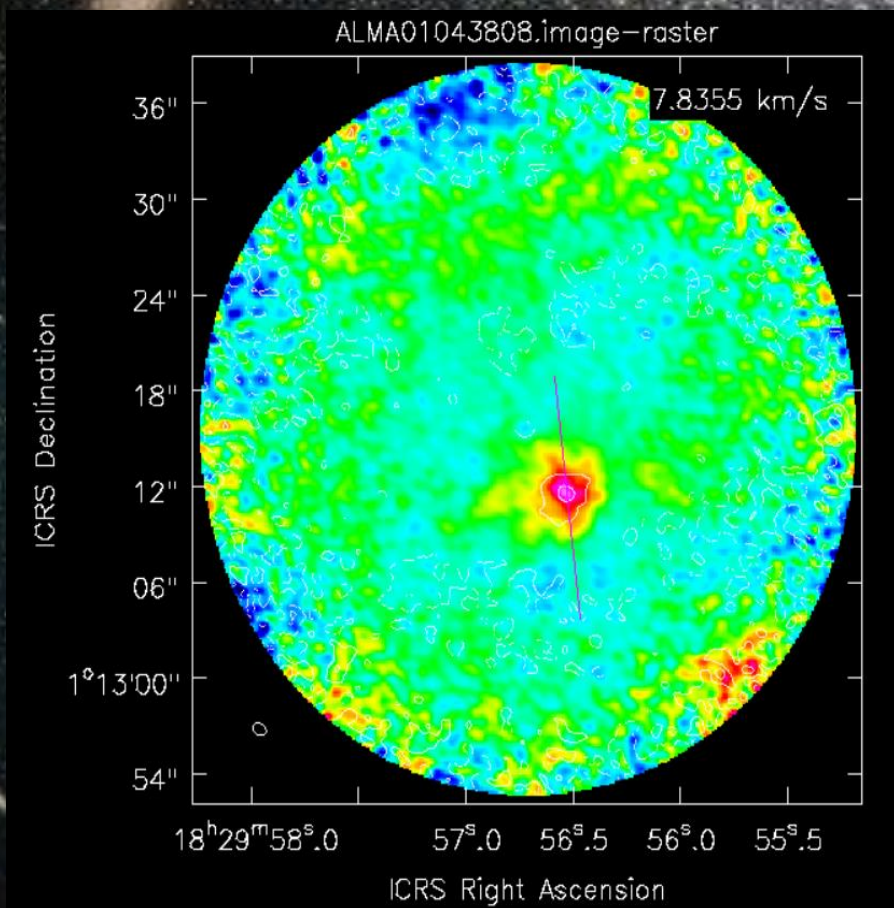


Pv diagram

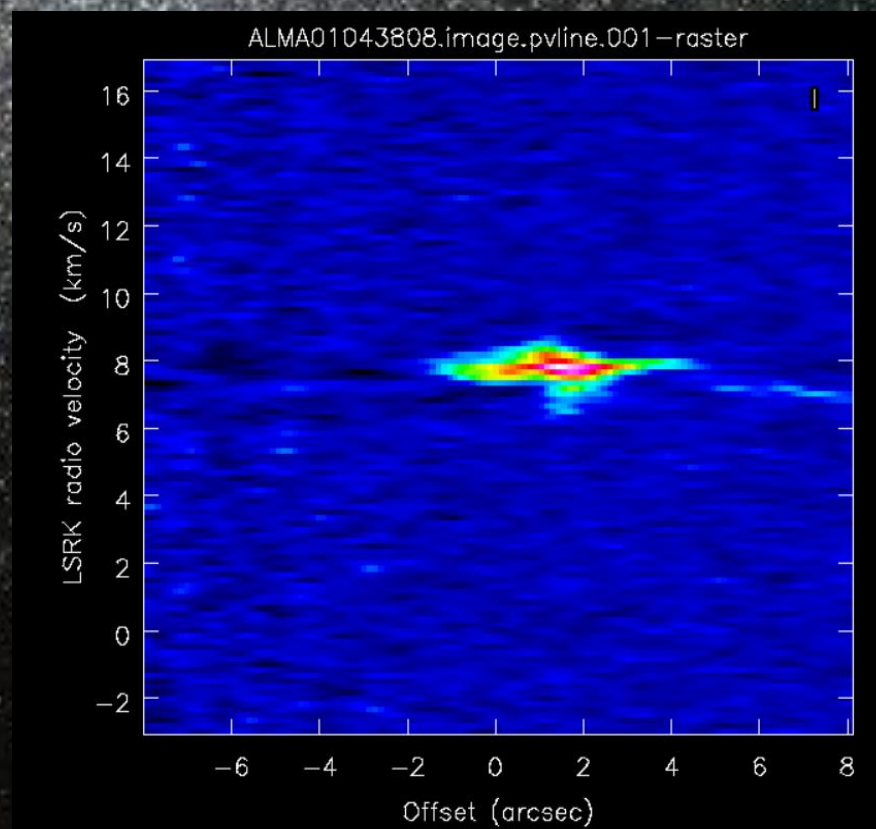
Pa=7



Pv diagram



Pa=98



Conclusion

source	Line	Keplerian disk	Mass of the disk
Lupus3 MMS (class 0)	C18O	△	×
serpens_smm4 (class 0)	C18O ,13CO	×	×
TMC1A (class 1)	C18O	√	0.019 solar mass
L1551NE (class 1)	C18O	×	0.046 solar mass
DG TAU B (class 1)	C18O	√	0.054 solar mass
L1448(class 0)	C18O,13CO	√	0.045 solar mass

reference

- ▶ <https://inspirehep.net/record/1242093/plots>
- ▶ <http://science.howstuffworks.com/how-are-stars-formed.htm>
- ▶ <https://arxiv.org/pdf/1411.0173.pdf>
- ▶ <http://iopscience.iop.org/article/10.3847/1538-4357/aa6116/pdf>
- ▶ <http://iopscience.iop.org/article/10.3847/1538-4357/834/2/178/pdf>

Thanks for listening