



image comes from wiki

Are icy worlds in our Solar system habitable?

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Outline:

1. Our goal
2. Introduction of Callisto
3. Observations
4. Approaches to find chemical composition
5. Results
6. Discussions
7. Summary

Our goal :

Using ALMA data to study chemical composition of Callisto to discuss whether they are habitable.



Io

Europa

Ganymede

Callisto

Introduction (Callisto):



image comes from wiki

1. Alternative name: Jupiter IV
2. Orbital distance to Jupiter: 1.88×10^6 km
3. Mean radius: 2410.3 km (0.378 Earths)
4. Mass: 1.076×10^{23} kg (0.018 Earths)
5. Mean density: 1.834 g/cm^3
6. Surface pressure: 7.4019×10^{-12} atm
7. Atmosphere: CO_2 & O_2

Observations:

- Atacama Large Millimetre/Submillimeter Array(ALMA)
- Frequency: 344.760 GHz~347.345 GHz, band 7
- Integration time: 206 seconds
- Callisto was observed as a bandpass calibrator
- Number of antennae: 21

Approaches to identify molecular transitions on the atmosphere of Callisto:

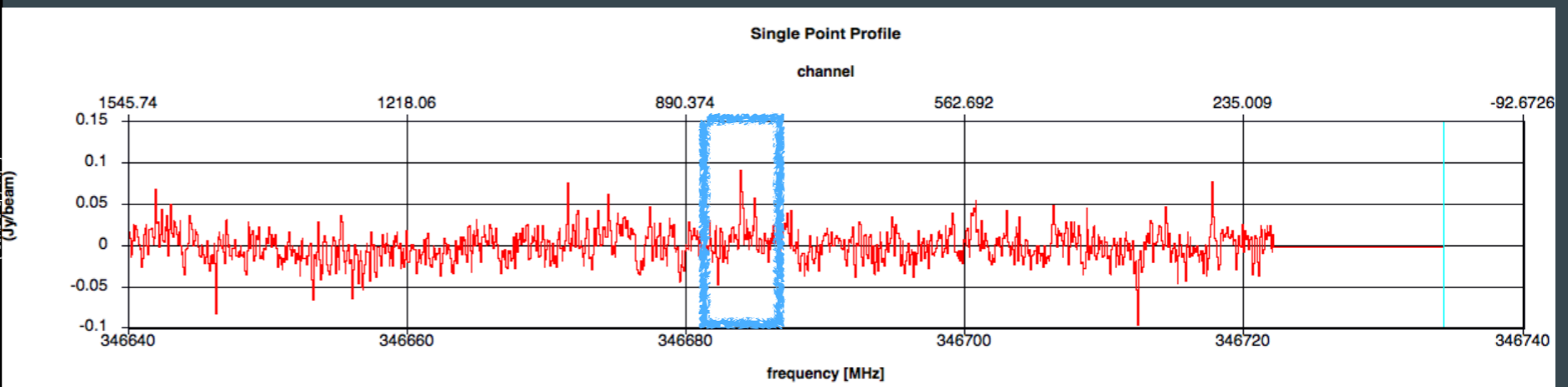
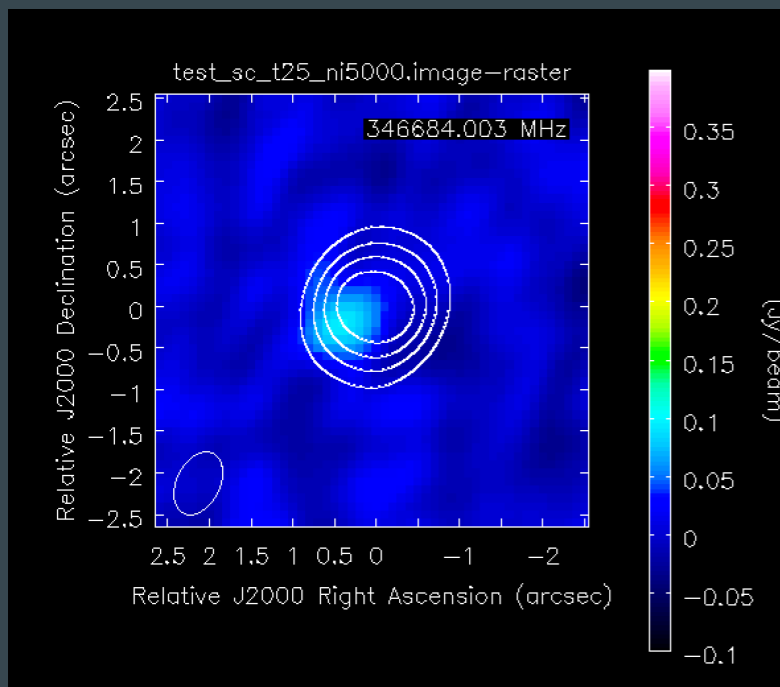
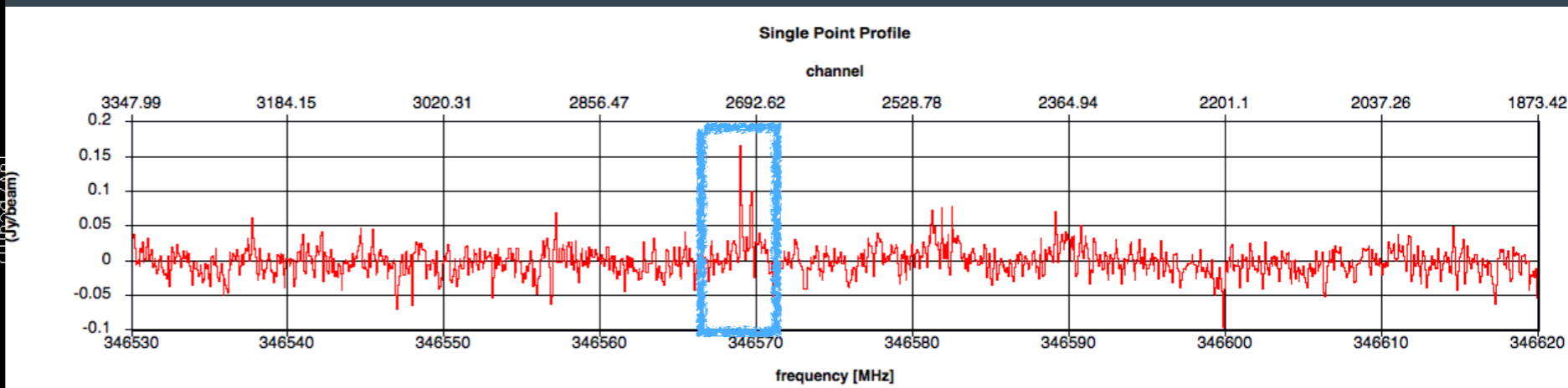
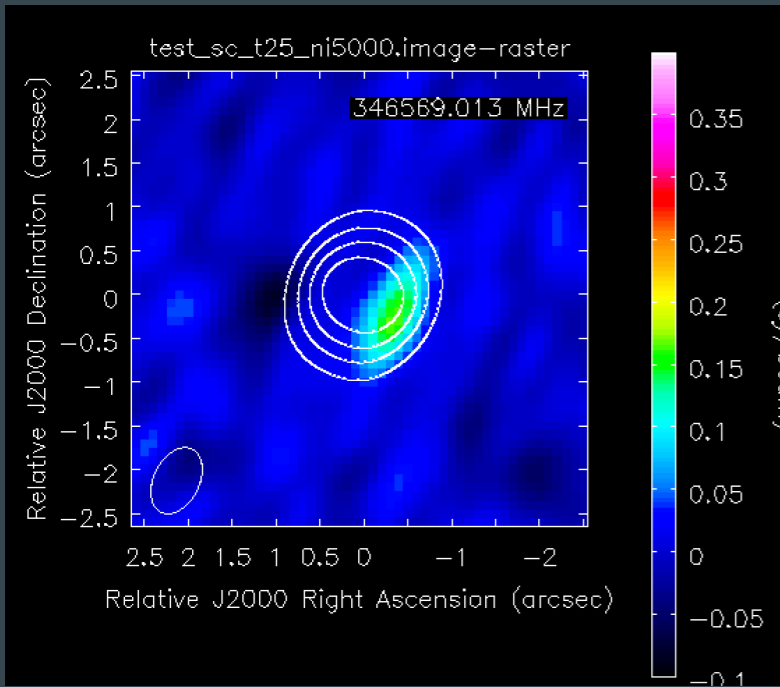
1. From spectral lines to molecular database
2. From molecular database to spectral lines

Approaches to identify molecular transitions on the atmosphere of Callisto:

1. From spectral lines to molecular database

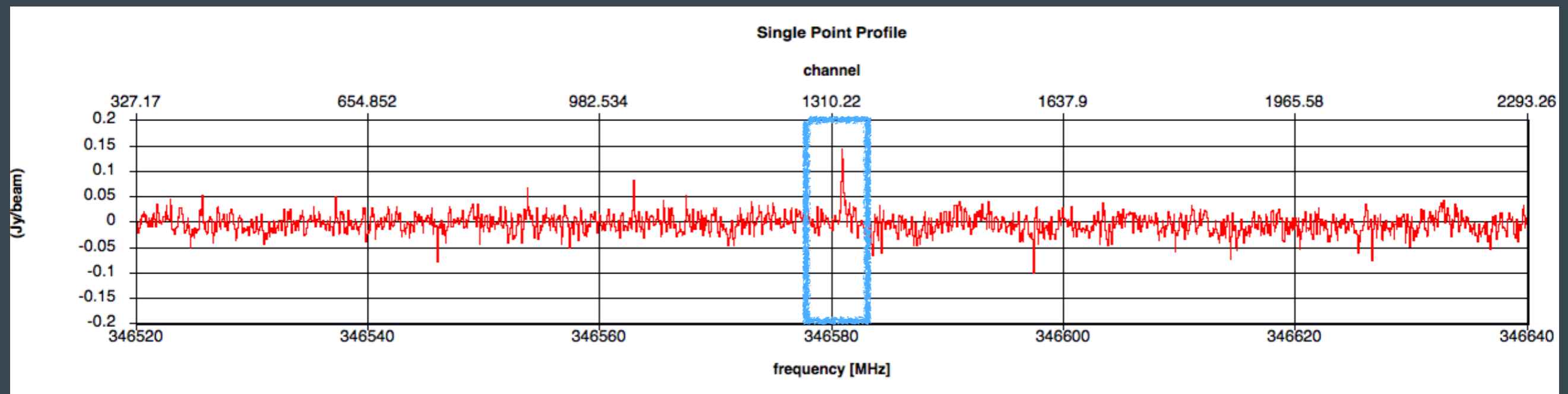
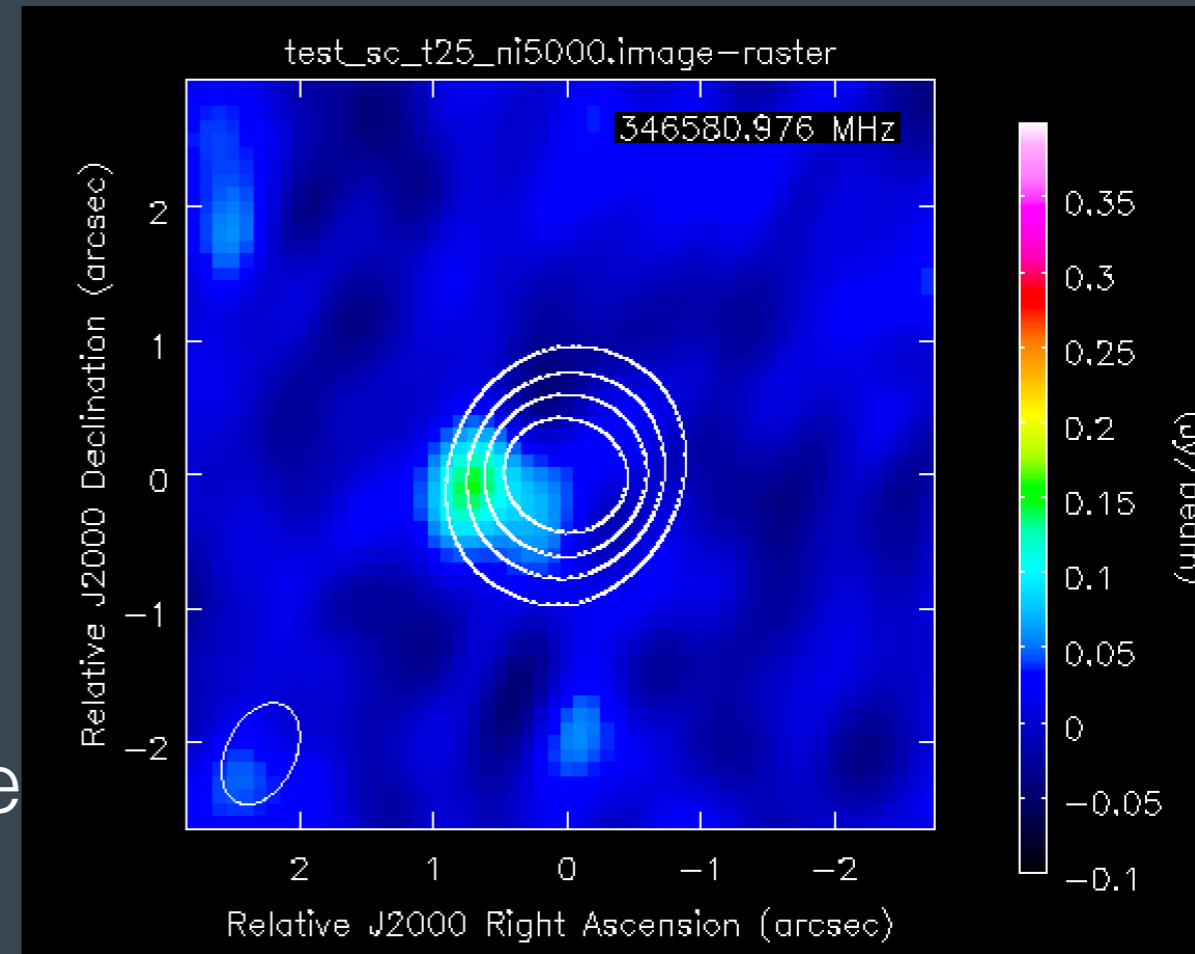
2. From molecular database to spectral lines

there are 22 candidate lines in spw0 (347111~347345 MHz)
 19 candidate lines in spw1 (344760~344994 MHz)
 16 candidate lines in spw2 (346550~346734 MHz)
 21 candidate lines in spw3 (345735~345969 MHz)



Selection criteria:

1. Intensity of the spectral features should be at least 3 sigmas.
2. Emission features should occur on Callisto disk.
3. There should not be strong noise peaks around Callisto.



However, Doppler shift should be considered...



Jet Propulsion Laboratory
California Institute of Technology

```
*****
Ephemeris / WWW_USER Tue Jul 26 02:20:49 2016 Pasadena, USA / Horizons
*****
Target body name: Callisto (504) {source: jup310}
Center body name: Earth (399) {source: DE431mx}
Center-site name: GEOCENTRIC
*****
Date__(UT)__HR:MN          r          rdot          delta          deldot          S-O-T /r          S-T-O          PsAng          PsAMV          GlxLon          GlxLat
*****
2012-Oct-17 09:33          5.052722997643  -1.1149049  4.37408152108482 -20.9213458  128.3039 /L          8.9096          n.a.          n.a.          179.996304  -12.615693
*****
```

$$\Delta\nu = \nu_{sky} - \nu_{rest} \quad \text{—————(1)}$$

$$\Delta\nu = \left(-\frac{V}{c}\right) \nu_{sky} \quad \text{—————(2)}$$

$\Delta\nu$: shifted frequency

V : velocity

for example:

Freq (sky): 346580.976 MHz

Freq (rest): 346556.997 MHz



Molecular Spectroscopy

Jet Propulsion Laboratory

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JPL Catalog Search Form

You need a Browser with Forms Capability to use this.

See [README](#) for output format.

What is the **minimum** frequency ?

What is the **maximum** frequency ?

The frequency units can be GHz or wavenumbers. If GHz is checked, the format of the output will be in standard catalog form (with MHz units). If wavenumber is checked, the frequency and error fields of the output will be in wavenumbers.

What is the common log of the **minimum** strength in catalog units ?

All

- 1001 H-atom
- 2001 D-atom
- 3001 HD
- 4001 H2D+
- 7001 Li-6-H

What molecules should be included ? (use mouse control click to select multiple values)

Response will be limited to 1000 lines. Press this button to submit the query: .

To reset the form, press this button: .

take 346580.976 MHz (sky freq) as an example:

	44006	DNCO									
346555.3361	0.0366	-5.0983	3	92.4976	33	44006	30417	01716	16	01616	
346556.1689	0.0353	-2.6133	3	92.4975	37	44006	30417	01718	16	01617	
346556.1706	0.0353	-2.6389	3	92.4976	35	44006	30417	01717	16	01616	
346556.1717	0.0353	-2.6645	3	92.4975	33	44006	30417	01716	16	01615	
346556.9570	0.0365	-5.0983	3	92.4975	35	44006	30417	01717	16	01617	
	46004	C2H5OH									
346550.6162	0.0147	-7.4278	3	752.7514	95	46004	140447	938 1	461037	1	
346552.5552	0.0093	-6.7288	3	633.4758	87	46004	140443	835 1	43 736	1	
	46006	NO2									
346558.0681	0.0143	-6.1981	3	107.4564	30	46006	33514	2121515	15	1151516	
	47002	HC-13-OOH									
346558.4520	0.0500	-2.7909	3	88.5244	31	-47002	30315	213	14	212	
	48007	O3-2v2									
346553.5307	0.2497	-7.7715	3	2071.1536	71	48007	140435	729 2	36	630	2
	48011	CH3OOH									
346558.4144	7.2242	-5.4326	3	535.6029138		48011	1404341223	1	341124	4	
346558.6272	7.2265	-5.4326	3	535.6030138		48011	1404341222	2	341123	5	
346558.7624	7.2248	-5.4326	3	535.6029138		48011	1404341223	0	341123	3	
346558.7661	7.2248	-5.4326	3	535.6029138		48011	1404341222	0	341124	3	
	49005	HO3									
346554.3318	0.7462	-5.9359	3	48.1212	25	49005	32511	3 91212	12	0121313	
346554.3683	0.7463	-5.9707	3	48.1212	23	49005	32511	3 91211	12	0121312	

take 346580.976 MHz (sky freq) as an example:

	44006	DNCO									
346555.3361	0.0366	-5.0983	3	92.4976	33	44006	30417	01716	16	01616	
346556.1689	0.0353	-2.6133	3	92.4975	37	44006	30417	01718	16	01617	
346556.1706	0.0353	-2.6389	3	92.4976	35	44006	30417	01717	16	01616	
346556.1717	0.0353	-2.6645	3	92.4975	33	44006	30417	01716	16	01615	
346556.9570	0.0365	-5.0983	3	92.4975	35	44006	30417	01717	16	01617	
	46004	C2H5OH									
346550.6162	0.0147	-7.4278	3	752.7514	95	46004	140447	938	1	461037	
346552.5552	0.0093	-6.7288	3	633.4758	87	46004	140443	835	1	43	
	46006	NO2									
346558.0681	0.0143	-6.1981	3	107.4564	30	46006	33514	2121515	15	1151516	
	47002	HC-13-OOH									
346558.4520	0.0500	-2.7909	3	88.5244							
	48007	O3-2v2									
346553.5307	0.2497	-7.7715	3	2071.1536							
	48011	CH3OOH									
346558.4144	7.2242	-5.4326	3	535.60291							
346558.6272	7.2265	-5.4326	3	535.6030138		48011	11404341222	2		341123	
346558.7624	7.2248	-5.4326	3	535.6029138		48011	11404341223	0		341123	
346558.7661	7.2248	-5.4326	3	535.6029138		48011	11404341222	0		341124	
	49005	HO3									
346554.3318	0.7462	-5.9359	3	48.1212	25	49005	32511	3	91212	12	
346554.3683	0.7463	-5.9707	3	48.1212	23	49005	32511	3	91211	12	

rest freq: 346556.997 MHz

Approaches to identify molecular transitions on the atmosphere of Callisto:

1. From spectral lines to molecular database
2. From molecular database to spectral lines



Molecular Spectroscopy

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California Institute of Technology

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What is the common log of the **minimum** strength in catalog units ?

What molecules should be included ? (use mouse control click to select multiple values)

46002 Si-30-O
46003 H2CS
46004 C2H5OH
46005 HCOOH
46006 NO2
46007 N2O-18

Response will be limited to 1000 lines. Press this button to submit the query: .

To reset the form, press this button: .

NO₂ transitions in frequency range 344-348 GHz

		46006	NO2								
345191.9942	0.0591	-6.4889	3	1101.7250102	46006	33550	2485051	49	3474950		
345196.8035	0.0590	-6.4972	3	1101.7277100	46006	33550	2485050	49	3474949		
345198.5321	0.0591	-6.5075	3	1101.7300 98	46006	33550	2485049	49	3474948		
345597.7244	0.0587	-6.4954	3	1101.6858100	46006	33550	2485150	49	3475049		
345602.4731	0.0585	-6.4874	3	1101.6884102	46006	33550	2485151	49	3475050		
345611.8231	0.0588	-6.4777	3	1101.6914104	46006	33550	2485152	49	3475051		
345642.4721	0.0123	-7.4677	3	107.4586 30	46006	33514	2121415	15	1151515		
345666.2468	0.0170	-7.4613	3	107.4605 28	46006	33514	2121414	15	1151514		
345710.0100	0.0600	-5.0605	3	107.4564 30	-46006	33514	2121415	15	1151516		
345721.5500	0.0600	-5.0898	3	107.4586 28	-46006	33514	2121414	15	1151515		
345753.9933	0.0252	-6.5615	3	1101.7300 94	46006	33548	4444847	49	3474948		
345754.8400	0.0600	-5.0900	3	107.4605 26	-46006	33514	2121413	15	1151514		
345766.3200	0.0600	-6.5591	3	1101.7277 96	-46006	33548	4444848	49	3474949		
345774.6600	0.0600	-6.5612	3	1101.7250 98	-46006	33548	4444849	49	3474950		
346357.3728	0.0198	-7.7046	3	107.4605 28	46006	33514	2121514	15	1151514		
346412.7036	0.0151	-6.2397	3	107.4586 28	46006	33514	2121514	15	1151515		
346460.0925	0.0149	-6.1899	3	107.4314 30	46006	33514	2121415	15	1151616		
346490.5947	0.0128	-7.5808	3	107.4586 30	46006	33514	2121515	15	1151515		
346510.7300	0.0600	-6.5613	3	1101.6884 98	-46006	33548	4444949	49	3475050		
346518.6100	0.0600	-6.5568	3	1101.6858 96	-46006	33548	4444948	49	3475049		
346531.7486	0.0263	-6.5329	3	1101.6914100	46006	33548	4444950	49	3475051		
346558.0681	0.0143	-6.1981	3	107.4564 30	46006	33514	2121515	15	1151516		
346606.8089	0.0163	-6.2302	3	107.4291 28	46006	33514	2121414	15	1151615		
346630.0547	0.0180	-7.6633	3	107.4564 32	46006	33514	2121516	15	1151516		
347287.5200	0.0600	-4.9976	3	107.4345 32	-46006	33514	2121516	15	1151617		
347297.9100	0.0600	-5.0836	3	107.4291 28	-46006	33514	2121514	15	1151615		
347308.1700	0.0600	-5.0563	3	107.4314 30	-46006	33514	2121515	15	1151616		
347375.8261	0.0126	-7.4940	3	107.4291 30	46006	33514	2121515	15	1151615		
347380.2017	0.0159	-7.4942	3	107.4314 32	46006	33514	2121516	15	1151616		
347809.0826	0.0163	-6.1647	3	252.3891 48	46006	33524	1232524	24	0242424		
347956.2481	0.0151	-5.6421	3	252.3866 50	46006	33524	1232525	24	0242425		

spw3

spw2

spw0

NO₂ transitions in frequency range 344-348 GHz

46006 NO2

*If NO₂ exist, a transition line is expected to be found at 347287MHz.

345753.9933	0.0252	-6.5615	3	1101.7300	94	46006	33548	4444847	49	3474948
345754.8400	0.0600	-5.0900	3	107.4605	26	-46006	33514	2121413	15	1151514
345766.3200	0.0600	-6.5591	3	1101.7277	96	-46006	33548	4444848	49	3474949
345774.6600	0.0600	-6.5612	3	1101.7250	98	-46006	33548	4444849	49	3474950

spw3

346510.7300	0.0600	-6.5613	3	1101.6884	98	-46006	33548	4444949	49	3475050
346518.6100	0.0600	-6.5568	3	1101.6858	96	-46006	33548	4444948	49	3475049
346531.7486	0.0263	-6.5329	3	1101.6914	100	46006	33548	4444950	49	3475051
346558.0681	0.0143	-6.1981	3	107.4564	30	46006	33514	2121515	15	1151516
346606.8089	0.0163	-6.2302	3	107.4291	28	46006	33514	2121414	15	1151615
346630.0547	0.0180	-7.6633	3	107.4564	32	46006	33514	2121516	15	1151516
347287.5200	0.0600	-4.9976	3	107.4345	32	-46006	33514	2121516	15	1151617
347297.9100	0.0600	-5.0836	3	107.4291	28	-46006	33514	2121514	15	1151615
347308.1700	0.0600	-5.0563	3	107.4314	30	-46006	33514	2121515	15	1151616

spw2

spw0

Results:

We identified 1 molecules

SO₂(Sulfur dioxide)

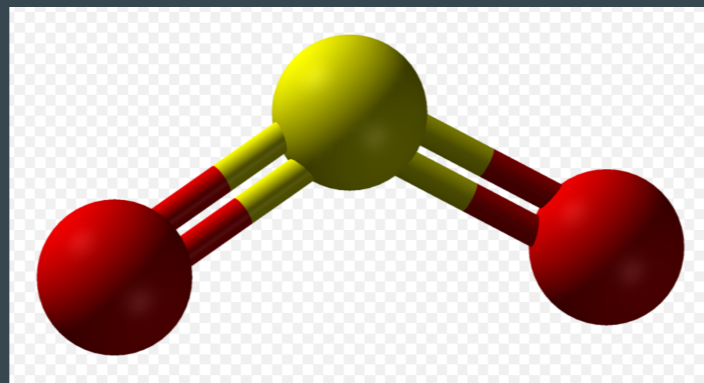


image comes from signaturewinelab

SO₂ transitions in frequency range 344-348 GHz

SO₂, v=0

345148.9708	0.0013	-3.9629	3	40.7151	11	64502	303	5	5	1	6	4	2
345338.5377	0.0015	-2.2398	3	53.1080	27	64502	30313	212			12	111	
345338.7778	0.0028	-8.7297	3	1669.2234	125	64502	303621646				631549		
345448.9841	0.0017	-3.6806	3	350.5898	53	64502	30326	917			27	820	
346523.8784	0.0013	-2.2075	3	102.7498	33	64502	30316	412			16	313	
346652.1691	0.0014	-1.9584	3	105.2994	39	64502	30319	119			18	018	
346910.9115	0.0049	-9.7834	3	1931.8670	135	64502	303671751				681652		
347829.2342	0.0038	-6.6123	3	1111.9644	115	64502	30357	652			56	749	

s1
s2

spw2

SO₂ line s1

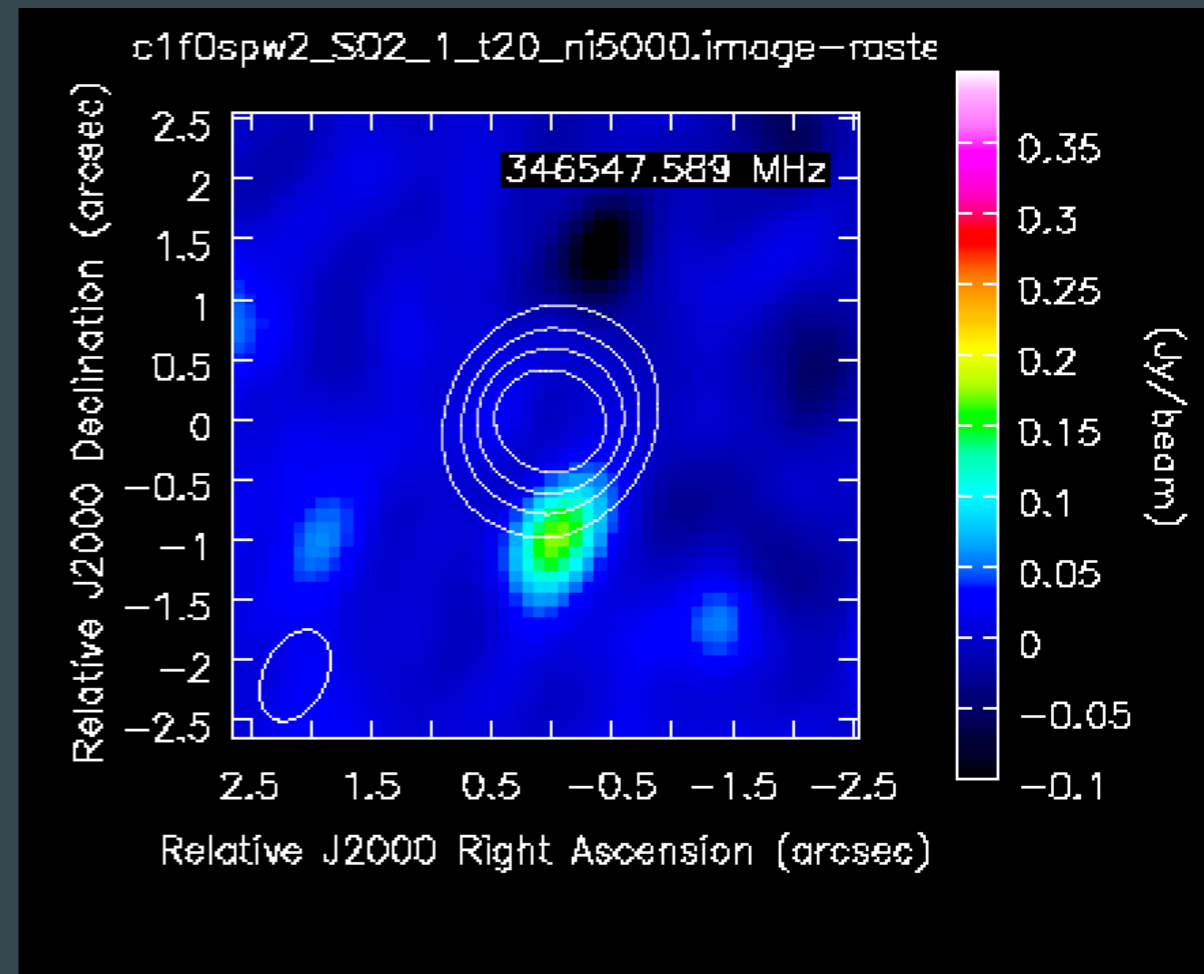
sky:346547.589 MHz

rest:346523.6121 MHz

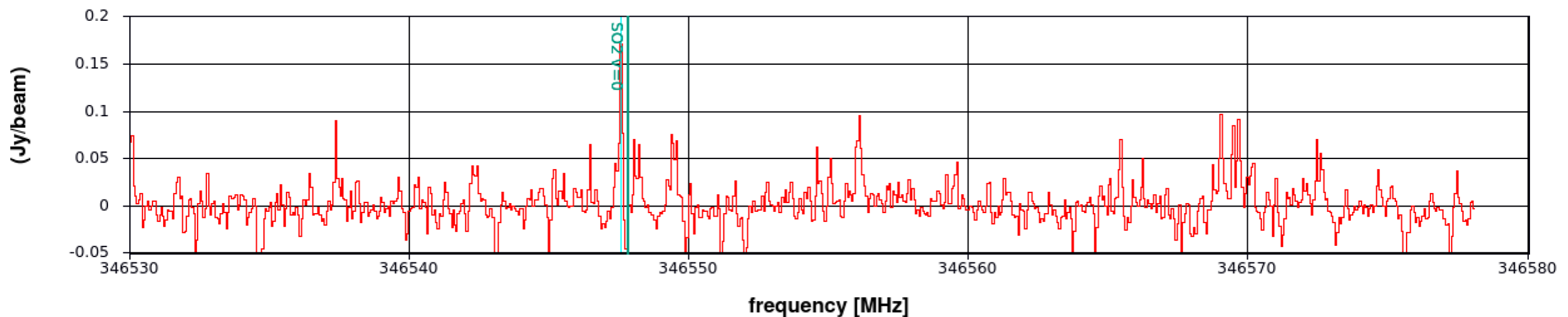
amp:173mJy/beam

SO₂, v=0

345148.9708	0.0013	-3.9629	3	40.7151	11	64502	303	5	5	1	6	4	2
345338.5377	0.0015	-2.2398	3	53.1080	27	64502	30313	212			12	111	
345338.7778	0.0028	-8.7297	3	1669.2234	125	64502	30362	1646			63	1549	
345448.9841	0.0017	-3.6806	3	350.5898	53	64502	30326	917			27	820	
346523.8784	0.0013	-2.2075	3	102.7498	33	64502	30316	412			16	313	
346652.1691	0.0014	-1.9584	3	105.2994	39	64502	30319	119			18	018	
346910.9115	0.0045	-9.7854	3	1951.8676	135	64502	30367	1751			68	1652	
347829.2342	0.0038	-6.6123	3	1111.9644	115	64502	30357	652			56	749	



Single Point Profile



SO₂ line s2

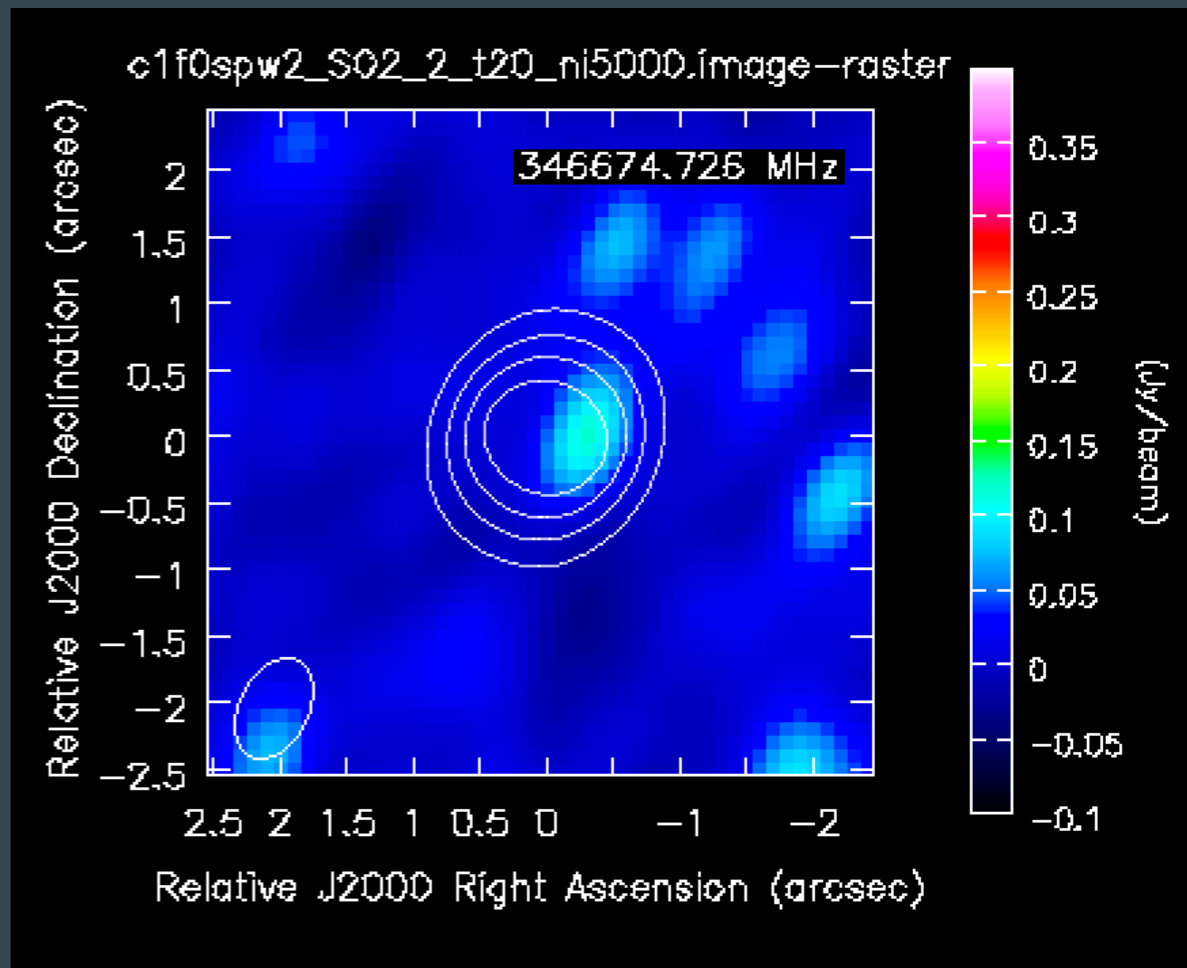
sky:346674.726 MHz

rest:346650.7403 MHz

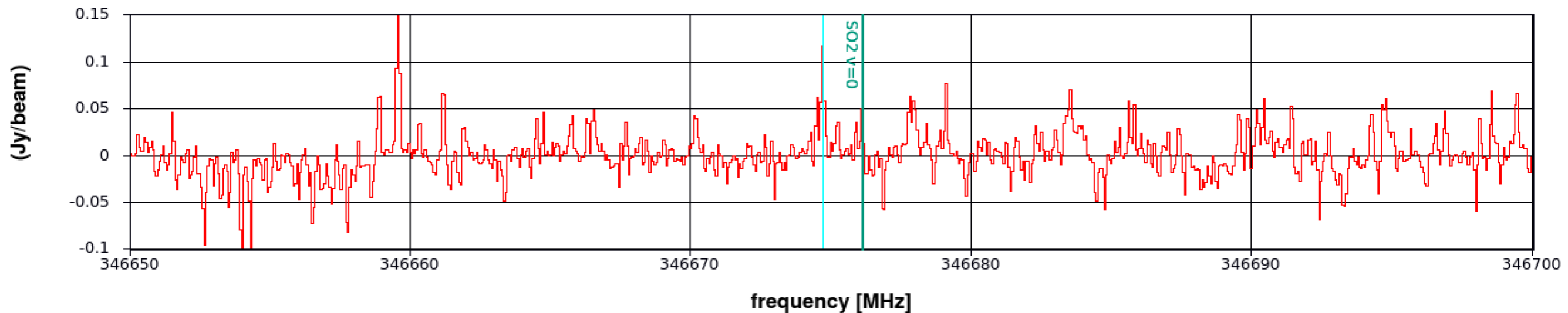
amp:118mJy/beam

SO₂, v=0

345148.9708	0.0013	-3.9629	3	40.7151	11	64502	303	5	5	1	6	4	2
345338.5377	0.0015	-2.2398	3	53.1080	27	64502	30313	212			12	111	
345338.7778	0.0028	-8.7297	3	1669.2234	125	64502	30362	1646			63	1549	
345448.9841	0.0017	-3.6806	3	350.5898	53	64502	30326	917			27	820	
346523.8784	0.0013	-2.2075	3	102.7498	33	64502	30316	412			16	313	
346652.1691	0.0014	-1.9584	3	105.2994	39	64502	30319	119			18	018	
346910.9115	0.0045	-5.7854	3	1951.8076	135	64502	30367	1751			68	1652	
347829.2342	0.0038	-6.6123	3	1111.9644	115	64502	30357	652			56	749	



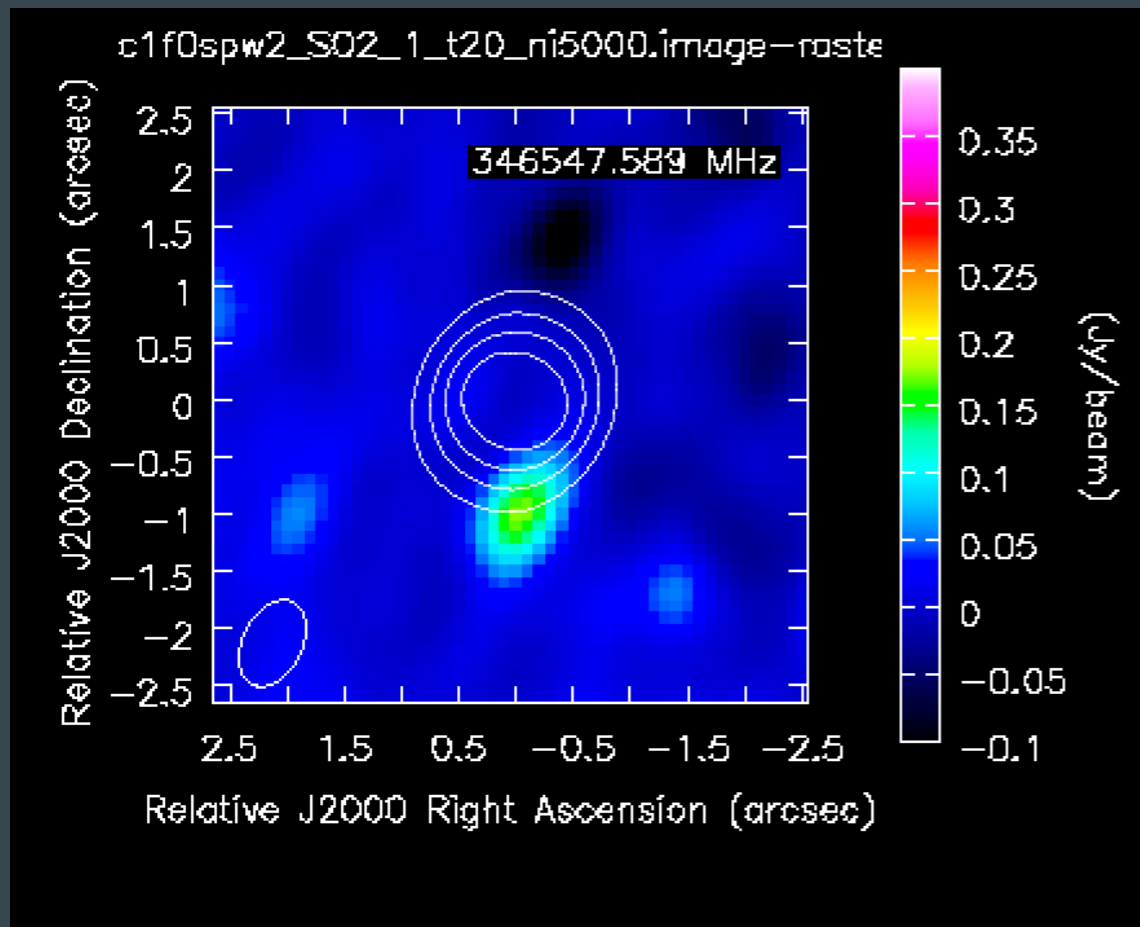
Single Point Profile



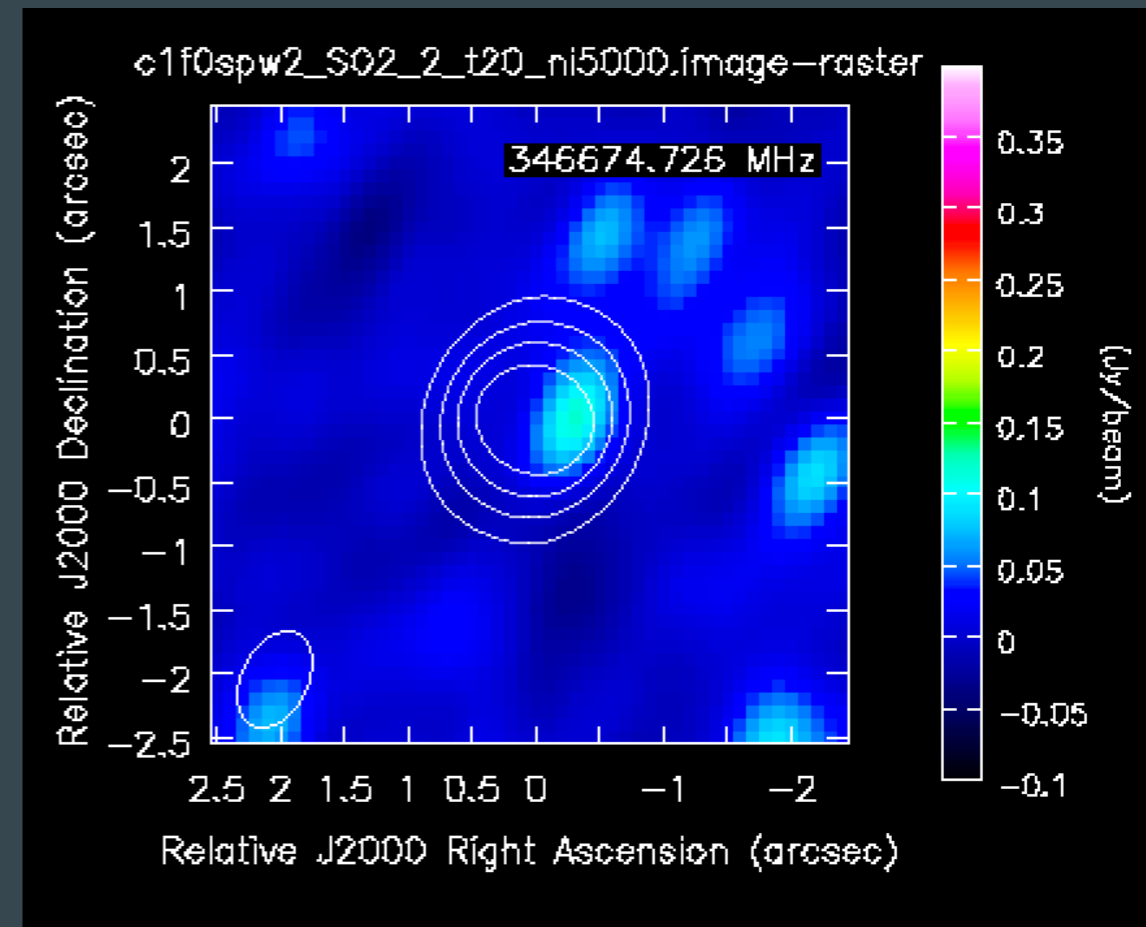
Discussions:

We have identified the transition line of SO_2 , however, there are still some questions need to be answered.

1. In the case of SO_2 , peaks don't show at same place on Callisto.



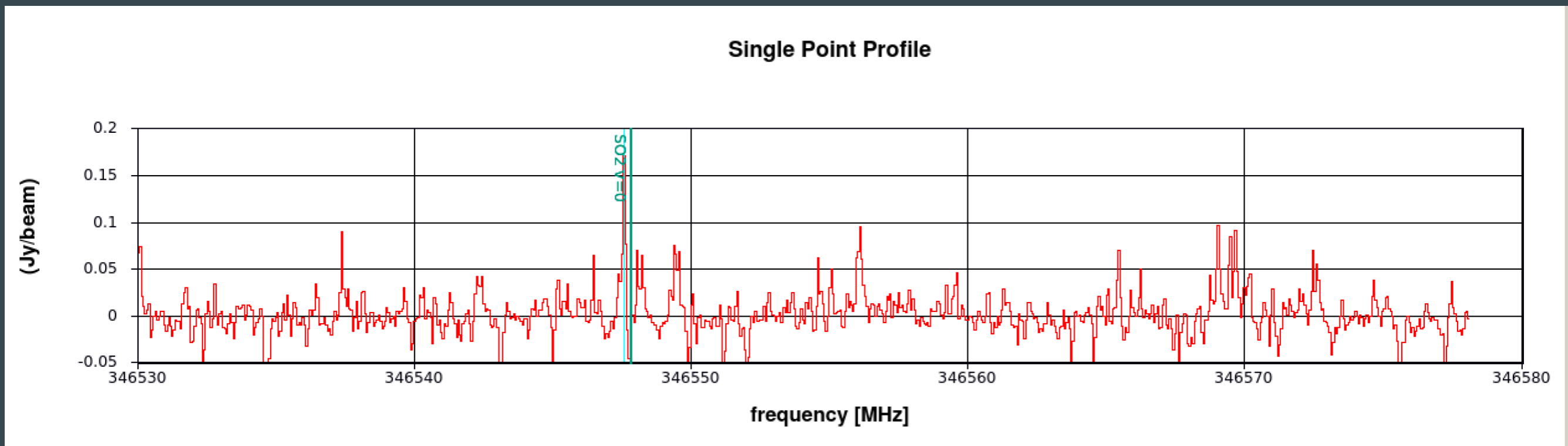
SO_2 line s1



SO_2 line s2

2. Observed frequency doesn't match the lab frequency precisely

Take SO₂ line s1 for example:



SO₂ line s1

Summary

1. High spectral and spatial resolution of ALMA helps us getting more information about Callisto.
2. We have found SO₂ on Callisto, however we still need more datas to prove it.
3. Finding SO₂ may help us know more about Callisto, especially for astrobiology.

The end